

# ML for statistically rigorous observing strategy optimization

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Alex I. Malz  
Carnegie Mellon University

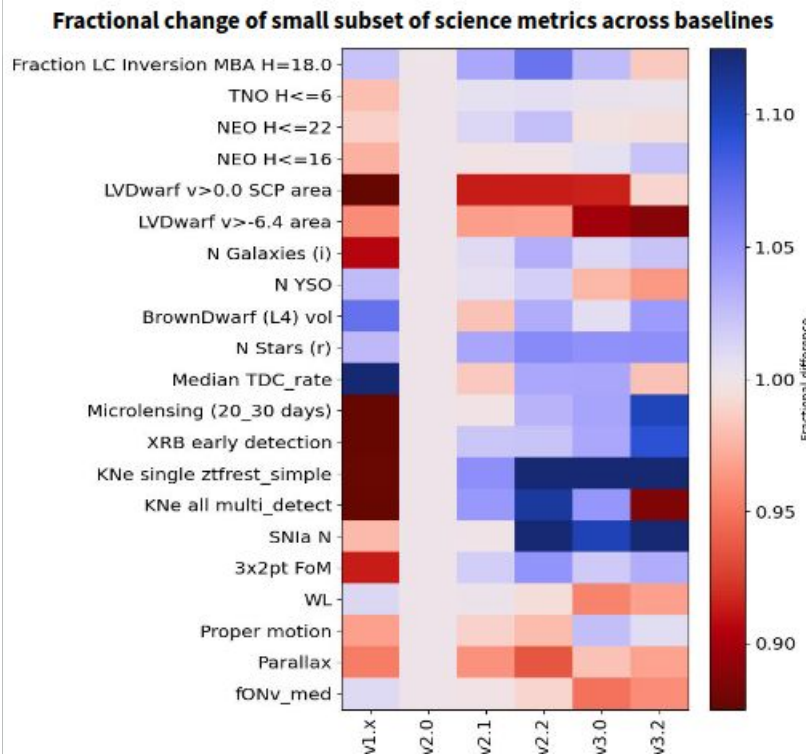


# Motivation: Observing strategy optimization

How can we choose the best possible observing strategy? (What does “best” even mean?)

- Every science case has its own metrics, with scaling properties in their own units.
- The choice is thus a convolution of nonuniform units and subjectivity of science goal importance.

**Can we put them all on common footing to isolate the subjective portion of the decision?**



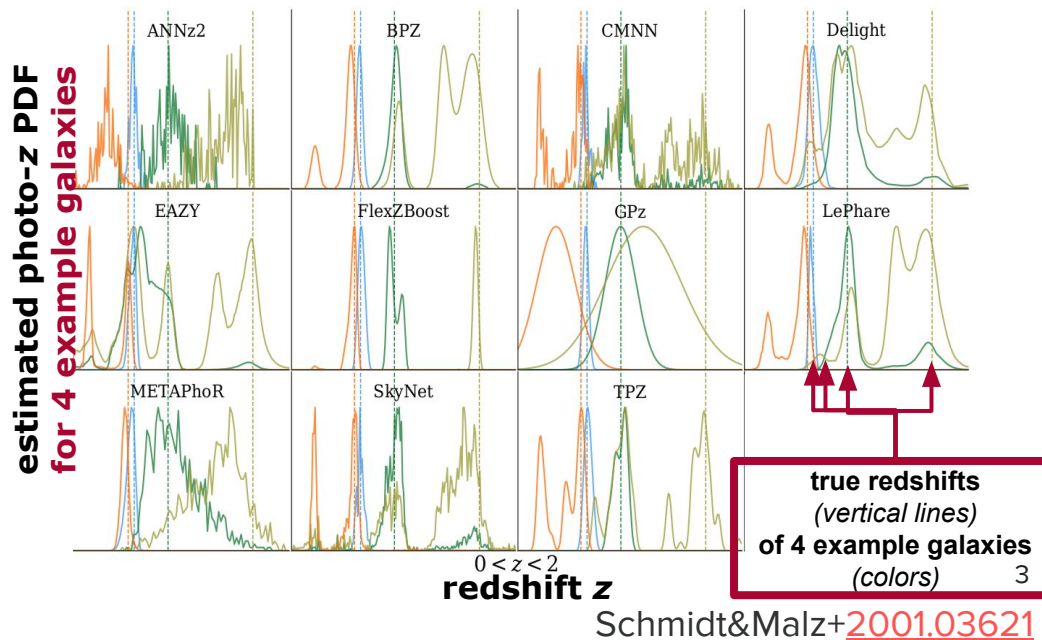
Credit: Lynne Jones

# Motivation: Photometric redshifts

Why can't we just add another metric for photo-z?

- There are many estimation algorithms to choose from, with no obvious forerunner.
- LSST's estimator (and prior) has not yet been chosen.

Can we derive a metric independent of the estimation algorithm?



Oh, and the implicit prior matters! See Malz&Hogg[2007.12178](#)

# What is TheLastMetric?

A science-agnostic information-theoretic metric,  $\mathfrak{n}$ , is the variational lower bound on the mutual information between redshift and photometry

$$I(Z; X_{phot}) \geq \mathbb{E}_{p(z, x_{phot})} [\log q_{\varphi}(z|x_{phot})] + H(Z) \equiv \mathfrak{n}$$

$$\mathcal{D}_{KL} [p(z|x_{phot})||q_{\varphi}(z|x_{phot})] + \mathbb{E}_{p(z, x_{phot})} [\log q_{\varphi}(z|x_{phot})] + H(Z)$$

$$\mathbb{E}_{p(z, x_{phot})} \left[ \log \frac{p(z|x_{phot})}{p(z)} \right]$$

$$H(Z) \equiv - \int dz p(z) \log p(z)$$

A variational approximation to  $\mathfrak{n}$  uses the **pzflow** conditional normalizing flow ([github.com/jfcrenshaw/pzflow](https://github.com/jfcrenshaw/pzflow))

with loss

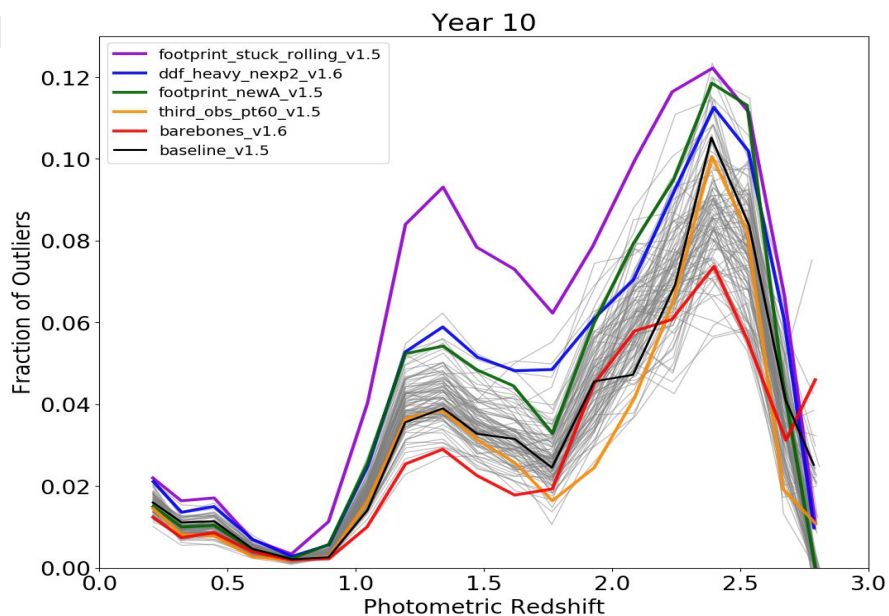
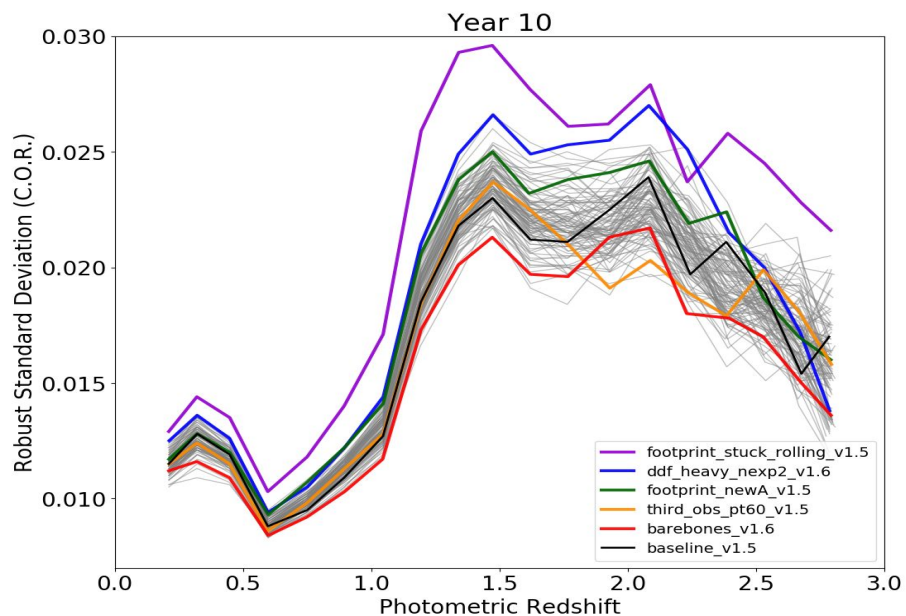
$$\mathcal{L} = -\mathbb{E}_{p(x, y)} [\log q_{\varphi}(x|y)]$$

for variational approximation

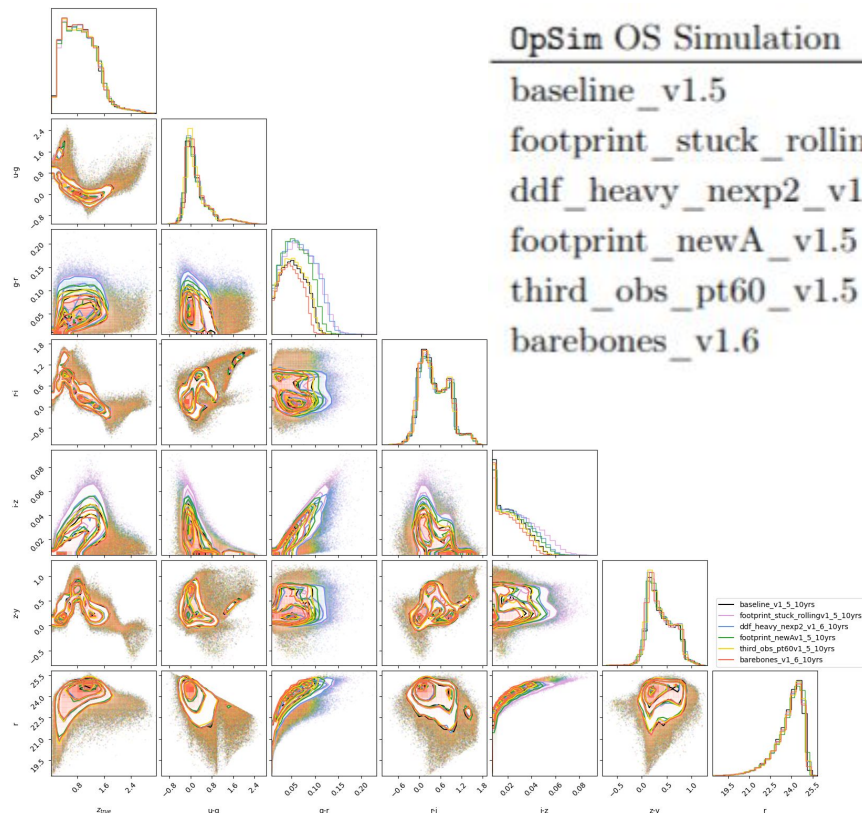
$$q_{\varphi}(x|y) = p(z = f_{\varphi}^{-1}(x; y)) \left| \det \frac{\partial f_{\varphi}}{\partial x}(x; y) \right|^{-1}$$

# Demonstration: goal

Explore the behavior of  $\mathfrak{n}$  relative to established metrics

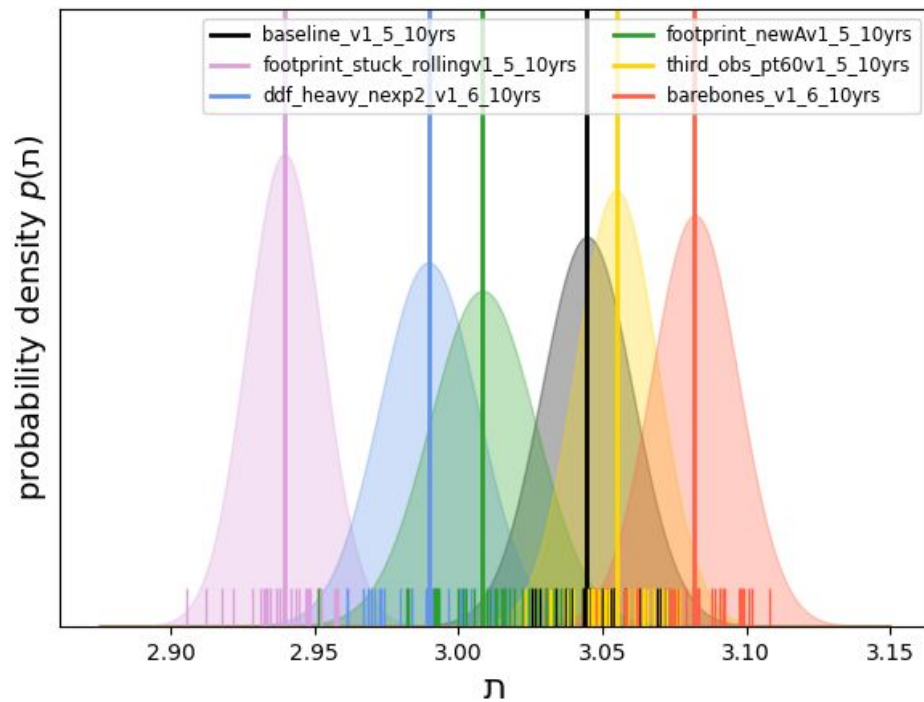
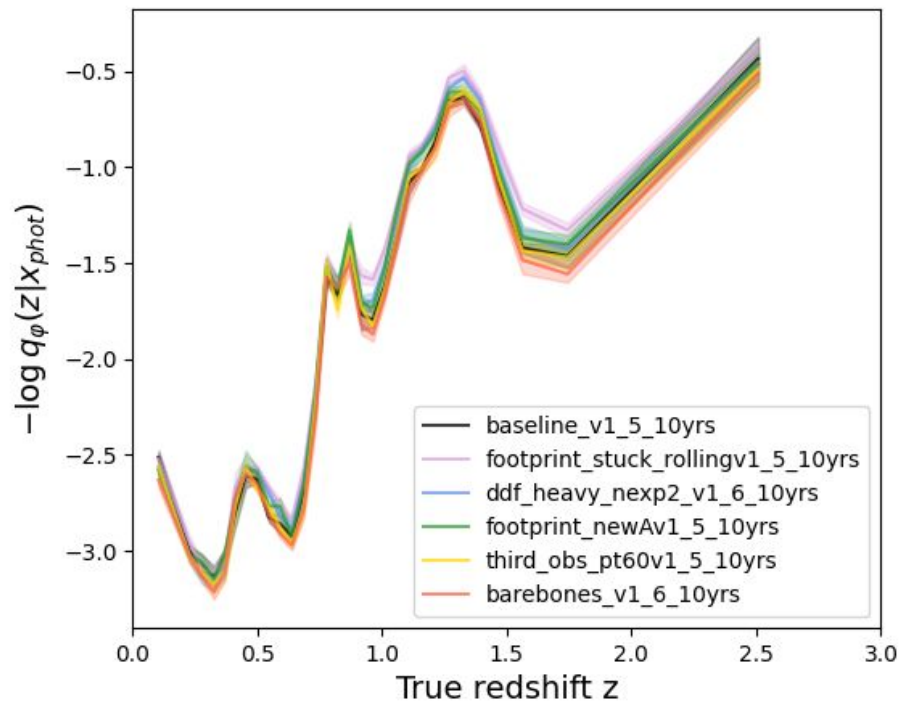


# Demonstration: experimental conditions

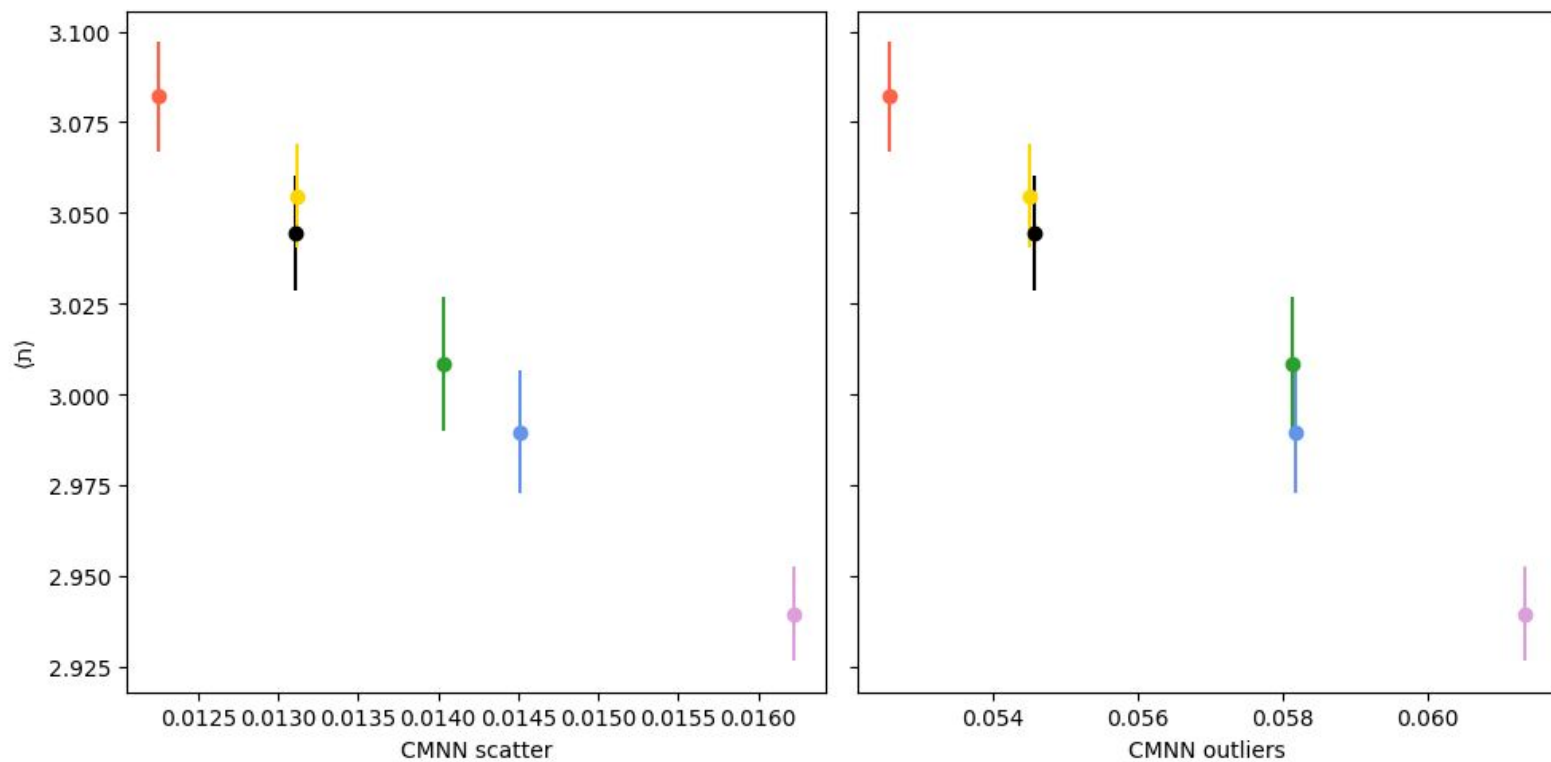


OpSim OS Simulation	m5_u	m5_g	m5_r	m5_i	m5_z	m5_y
baseline_v1.5	25.86	27.02	26.99	26.42	25.70	24.94
footprint_stuck_rolling_v1.5	25.56	26.68	26.62	26.06	25.33	24.61
ddf_heavy_nexp2_v1.6	25.57	26.82	26.84	26.26	25.57	24.82
footprint_newA_v1.5	25.75	26.87	26.85	26.29	25.55	24.78
third_obs_pt60_v1.5	25.87	27.03	26.99	26.43	25.70	24.93
barebones_v1.6	26.00	27.13	27.07	26.57	25.78	25.05

# Demonstration: results



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# Summary

Malz, Lanusse, Crenshaw, Scott & Graham (sub. to ApJS) [arXiv:2104.08229](https://arxiv.org/abs/2104.08229)

## Could $\pi$ really be “the last metric”?

- TheLastMetric quantifies the redshift information content of a photometric galaxy catalog.
- TheLastMetric distinguishes observing strategies without committing to any estimator (but it does have a prior).
- TheLastMetric has moderate computational cost, above the metrics of magnitude limits but below most photo-z estimators.

## Status of ongoing work

- Currently using TheLastMetric to perform comparisons of joint photometry between Rubin and other surveys (w/Bryan Scott)
- Beginning to implement TheLastMetric in MAF (w/Xiaolong Li)

## Next steps

- Adapt TheLastMetric to other scalar parameters derived from photometry.