

PLAN

Modelling dark matter dynamics

Phase space

Eulerian Perturbation Theory

Lagrangian Perturbation Theory

Spherical Collapse

Clustering statistics

Two-point statistics

Three-point statistics

One-point statistics + ...

SPHERICAL COLLAPSE

Newtonian force law

$$\ddot{R} = -\frac{GM(R)}{R^2} = -\frac{4\pi G}{3}\rho(t)R \quad \Big| \cdot 2\dot{R},$$

$$R^2 = -2GMR^{-1} - C$$

SPHERICAL COLLAPSE

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Parametric solution: cycloid

$$R = GM(1 - \cos \varphi)/C$$

$$t = GM(\varphi - \sin \varphi)/C^{3/2}.$$

SPHERICAL COLLAPSE

Parametric solution: cycloid

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translate to densities, mass conserved

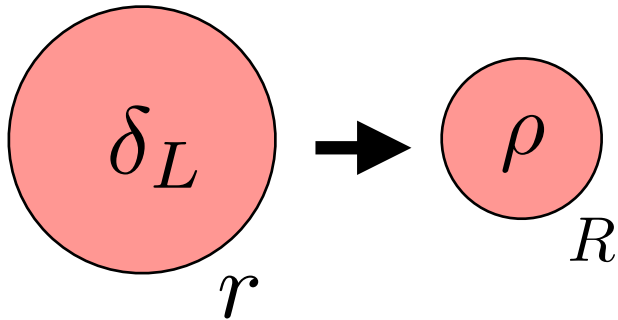
$$\delta_L = \frac{a(t)}{a_0} \propto \left(\frac{t_0}{t}\right)^{2/3} = \frac{3}{5} \left[\frac{3}{4}(\varphi - \sin \varphi) \right]^{2/3}$$

$$1 + \delta = \left(a(t) \frac{R_0}{R} \right)^3 \propto \left(\frac{t_0}{t}\right)^2 \left(\frac{R_0}{R}\right)^3 = \frac{9(\varphi - \sin \varphi)^2}{2(1 - \cos \varphi)^3}$$

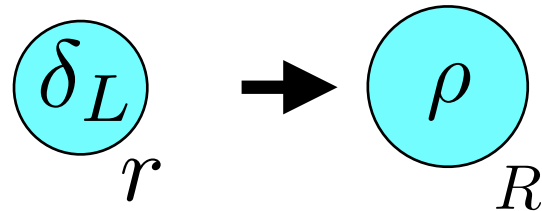
SPHERICAL COLLAPSE

mass conservation $r = \rho^{1/3} R$

overdensity



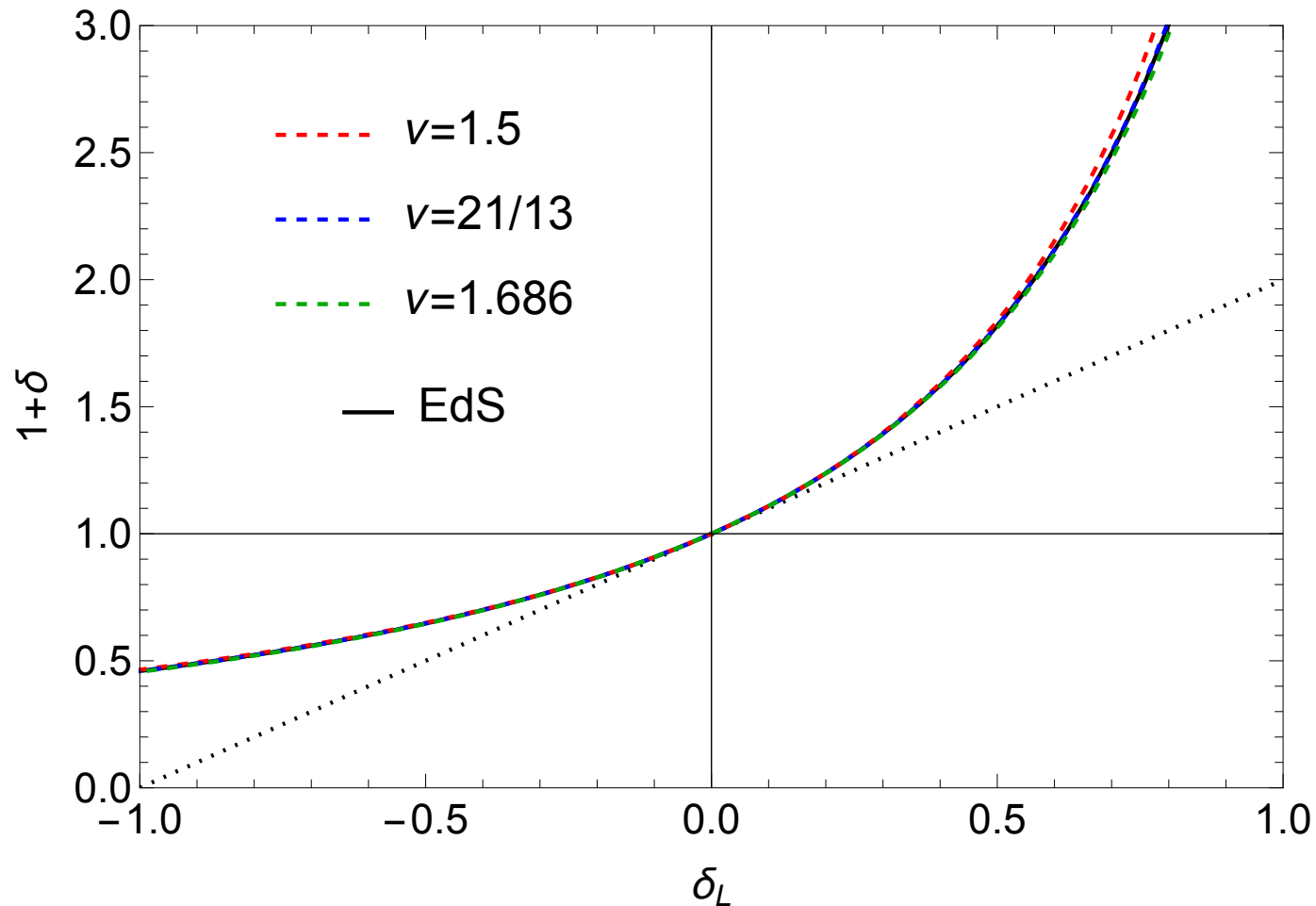
under density



SPHERICAL COLLAPSE

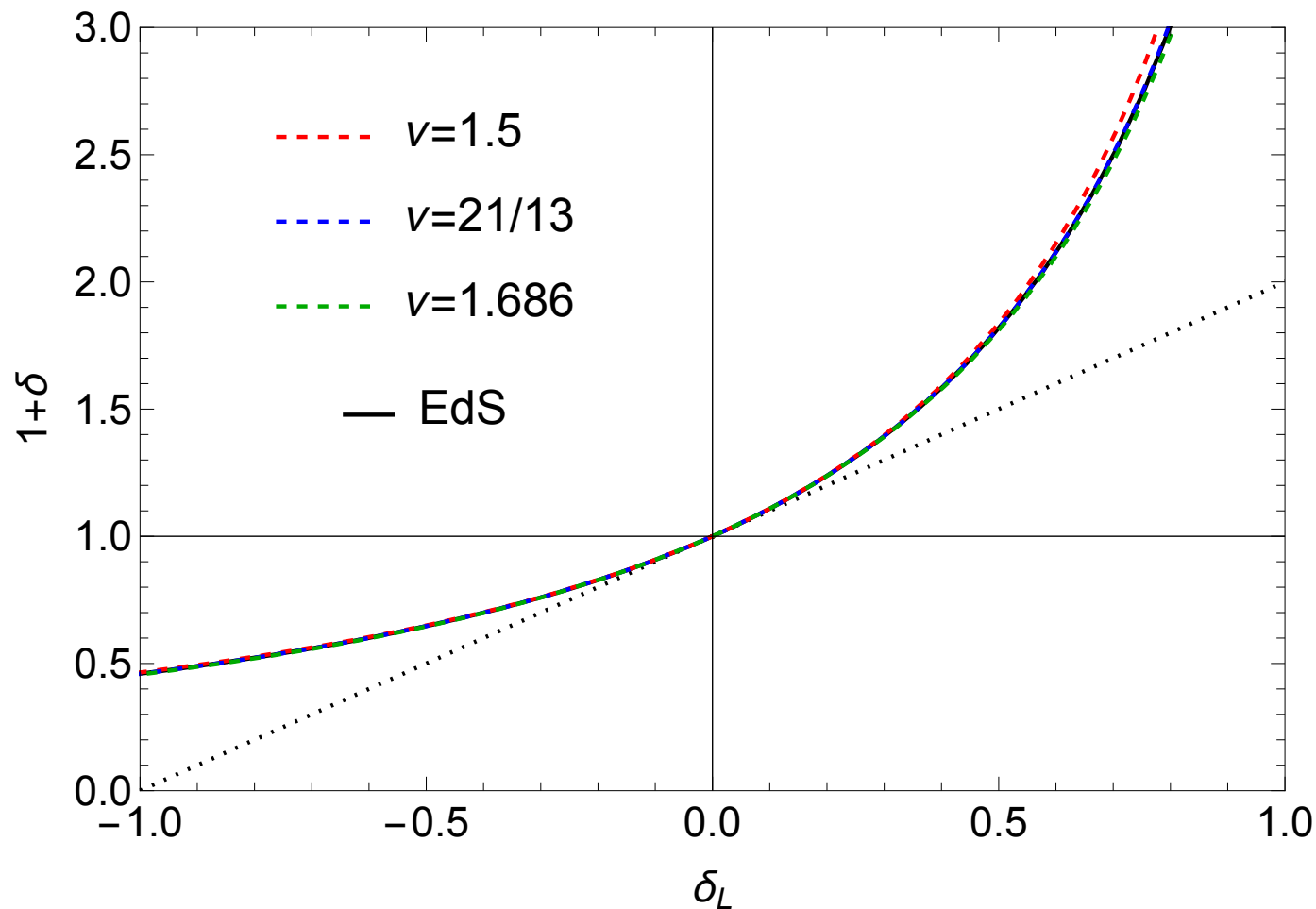
$$1 + \delta = \frac{9(\varphi - \sin \varphi)^2}{2(1 - \cos \varphi)^3}$$

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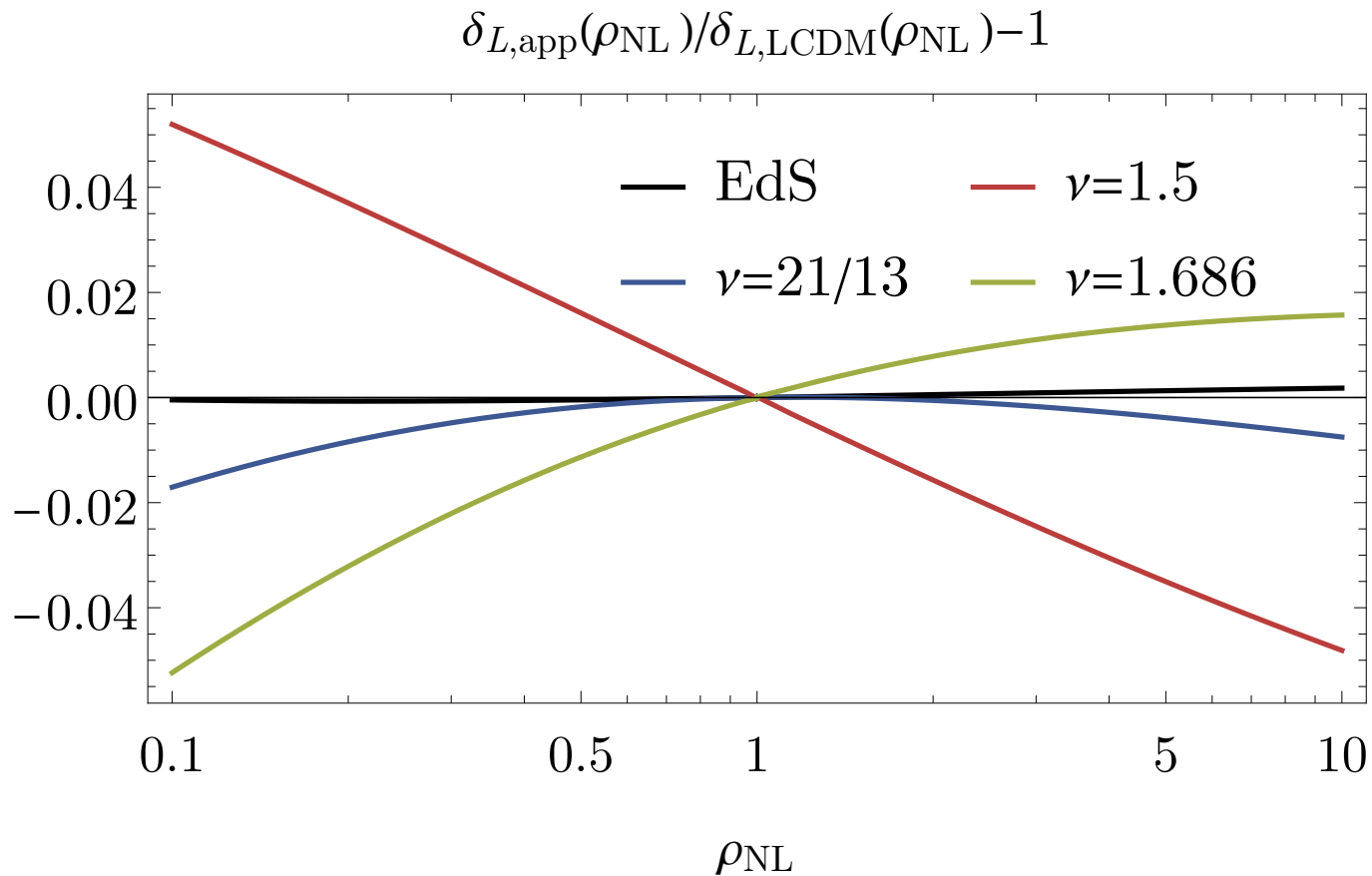
SPHERICAL COLLAPSE

$$1 + \delta = \left(1 - \frac{\delta_L}{\nu}\right)^{-\nu} \approx 1 + \delta_L + \frac{\nu + 1}{\nu} \frac{\delta_L^2}{2}$$



SPHERICAL COLLAPSE

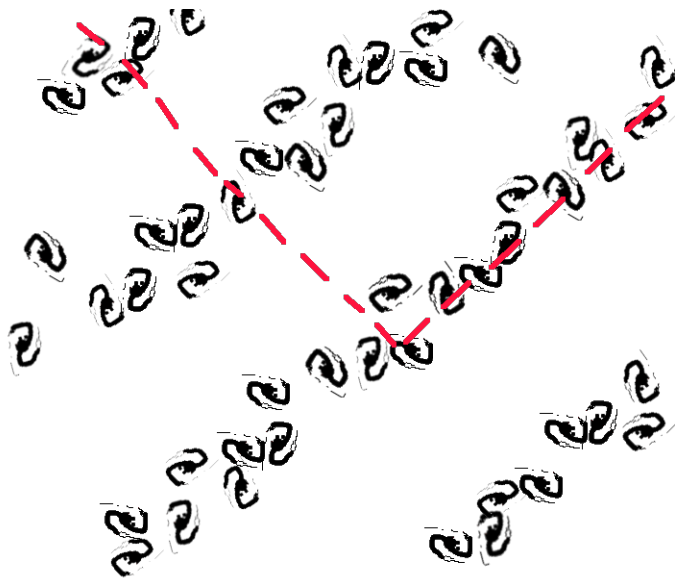
density mapping $\delta_L(\rho) \simeq \frac{21}{13} \left(1 - \rho^{-\frac{13}{21}} \right)$
~ cosmology independent



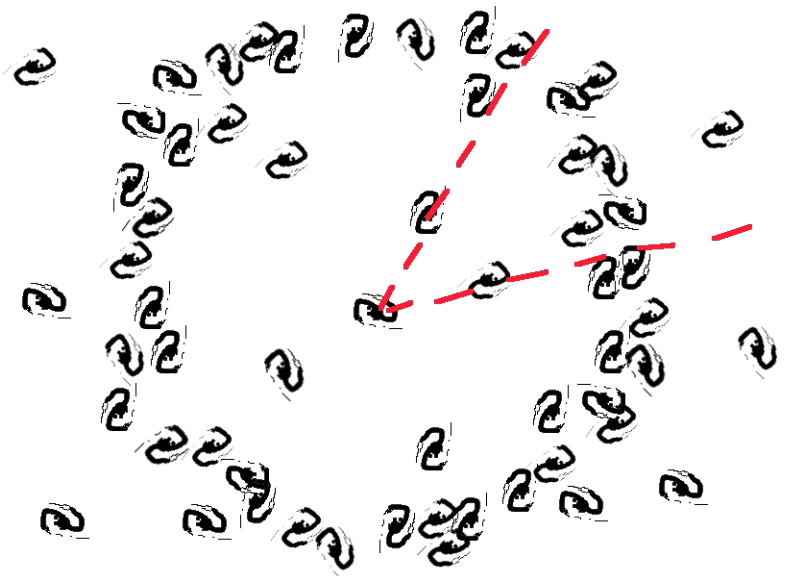
GALAXY CORRELATION FUNCTION

on large scales: matter distribution statistically

homogenous



isotropic

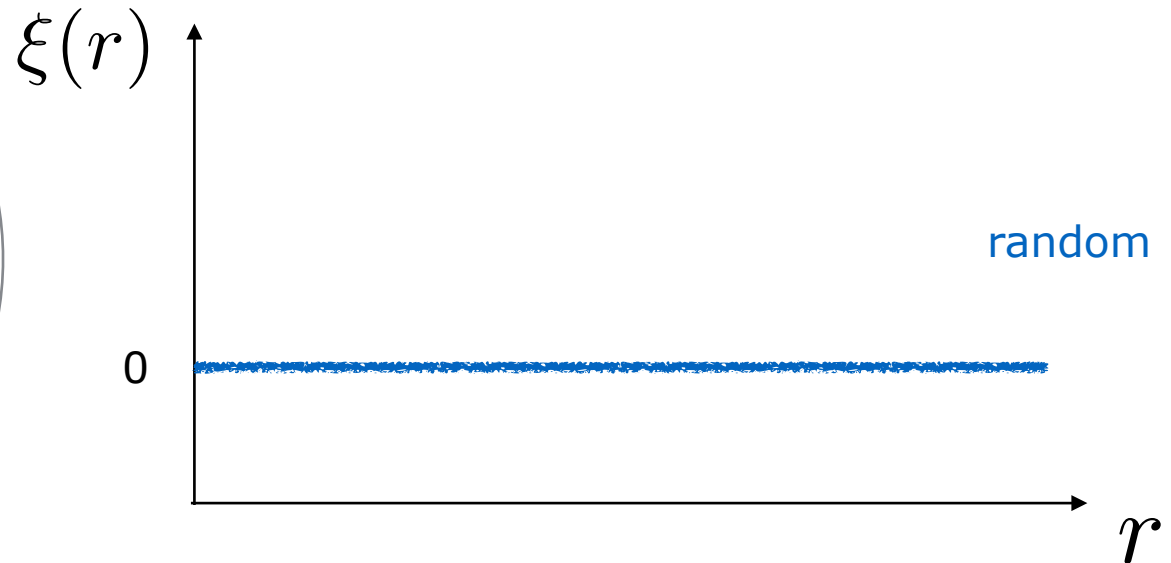
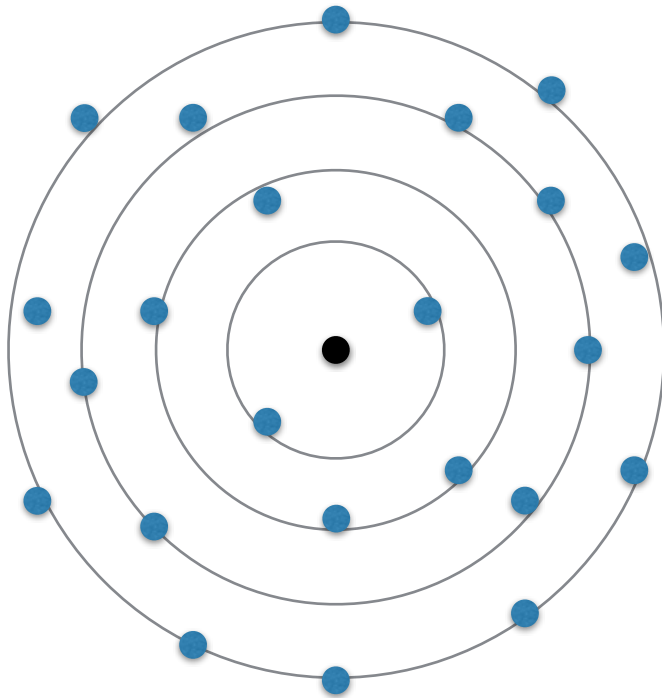


GALAXY CORRELATION FUNCTION

excess correlation compared to random distribution

$$\xi(r) = \langle \delta(\boldsymbol{x} + \boldsymbol{r}) \delta(\boldsymbol{x}) \rangle$$

↑ ↑
density contrast average

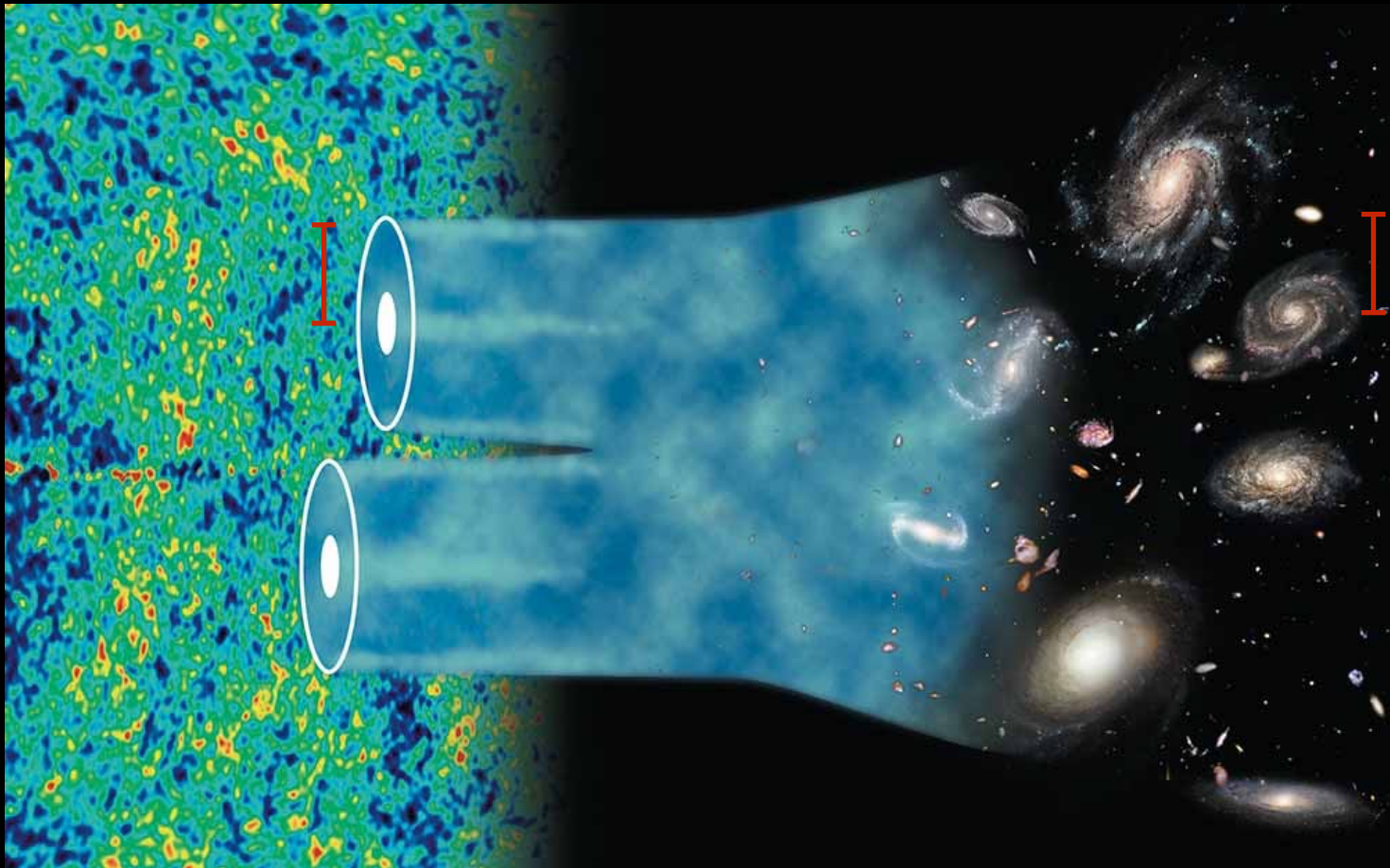


CMB $\rightarrow$ LARGE-SCALE STRUCTURE

Baryon Acoustic Oscillations survive all the way

Early Time
CMB

Late Time
Galaxy Distribution



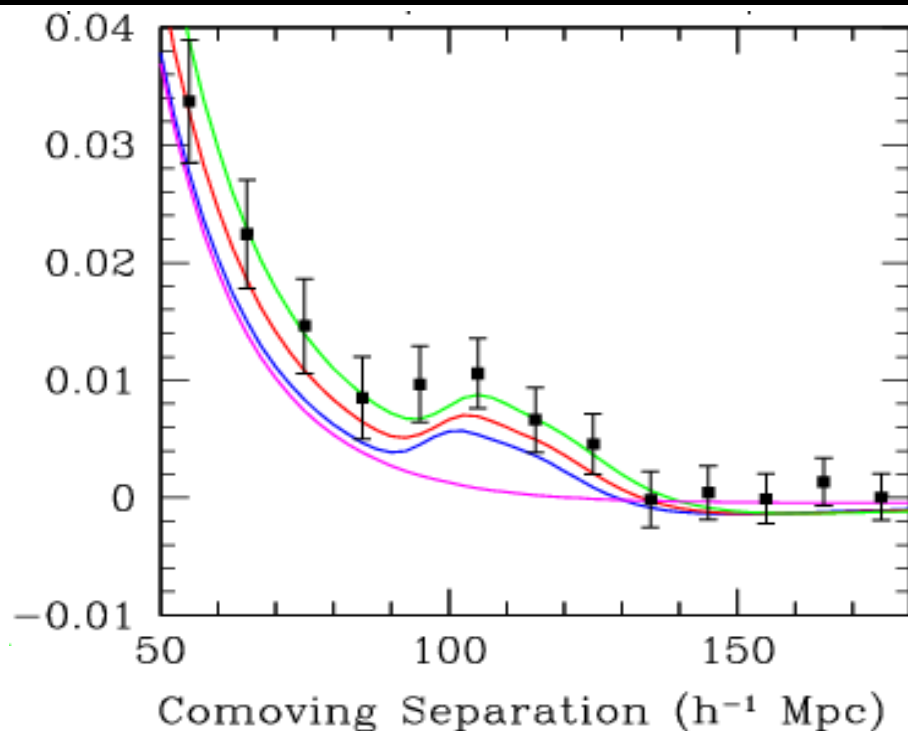
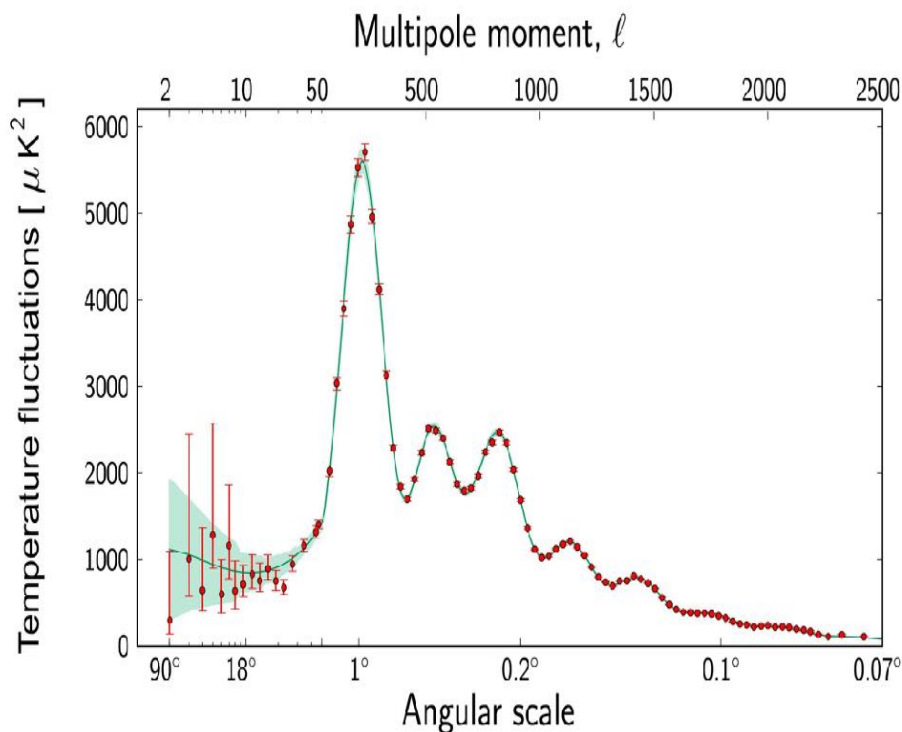
A COHERENT PICTURE

Baryon Acoustic Oscillations survive all the way

Early Time
CMB

Late Time
Galaxy Distribution

peak series in frequency scale = 1 peak in spatial scales

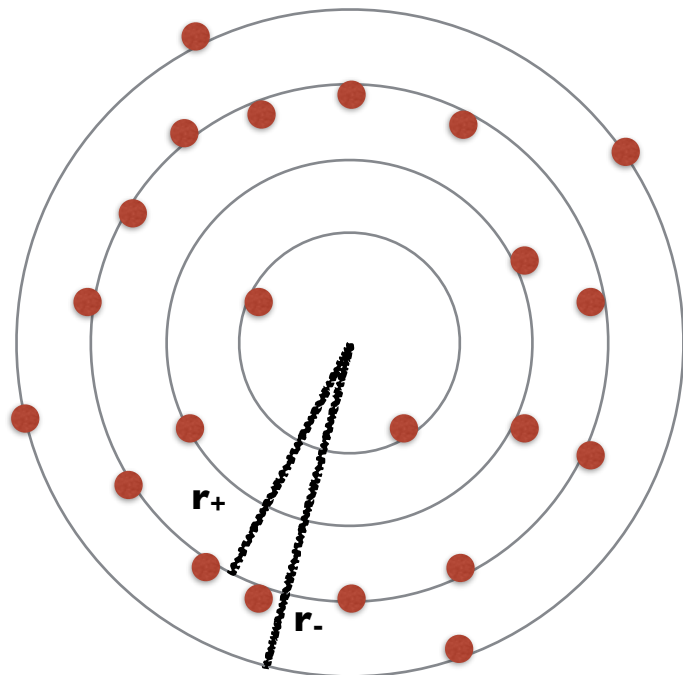


GALAXY CORRELATION FUNCTION

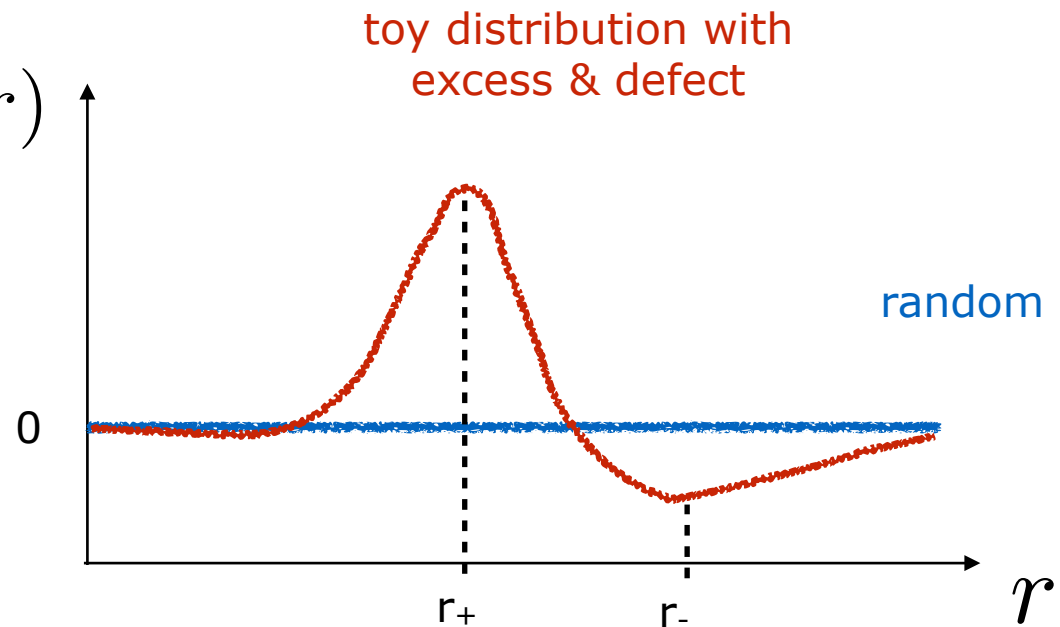
excess correlation compared to random distribution

$$\xi(r) = \langle \delta(x+r) \delta(x) \rangle$$

↑ ↑
density contrast average



$\xi(r)$

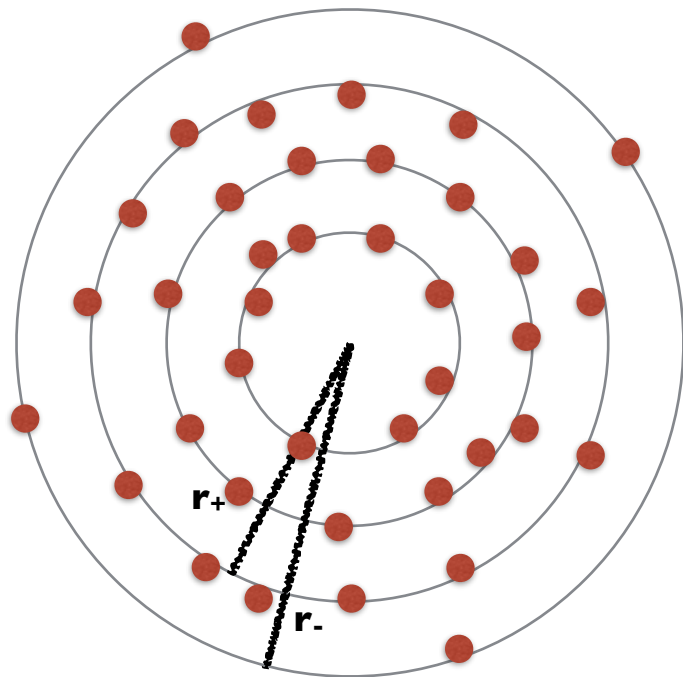


GALAXY CORRELATION FUNCTION

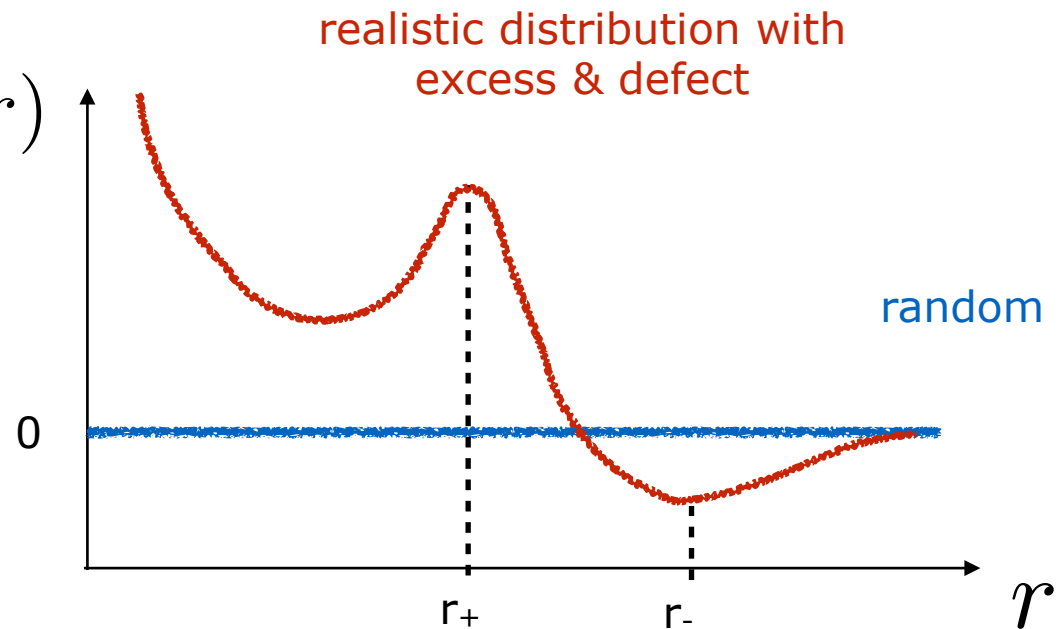
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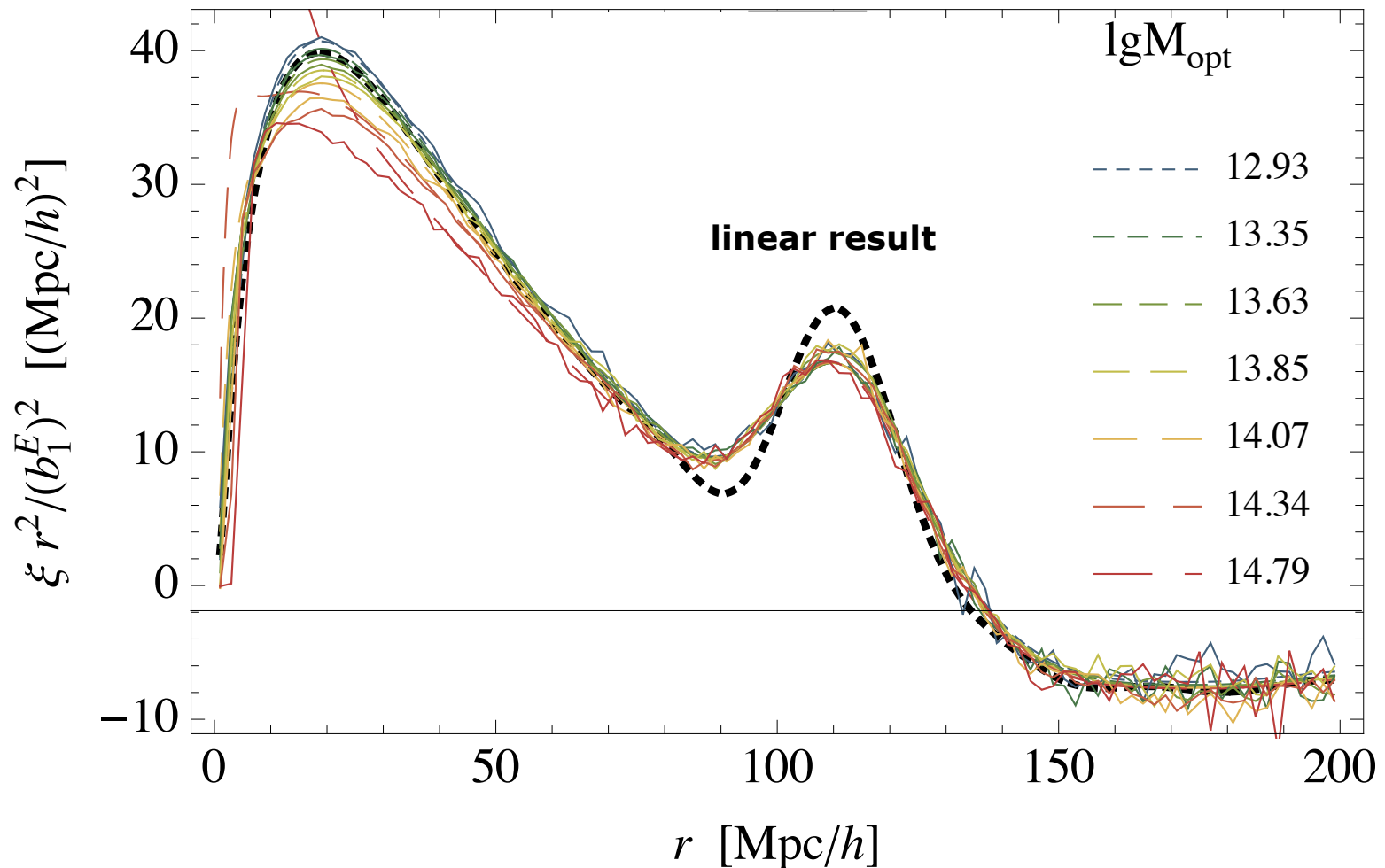


$\xi(r)$



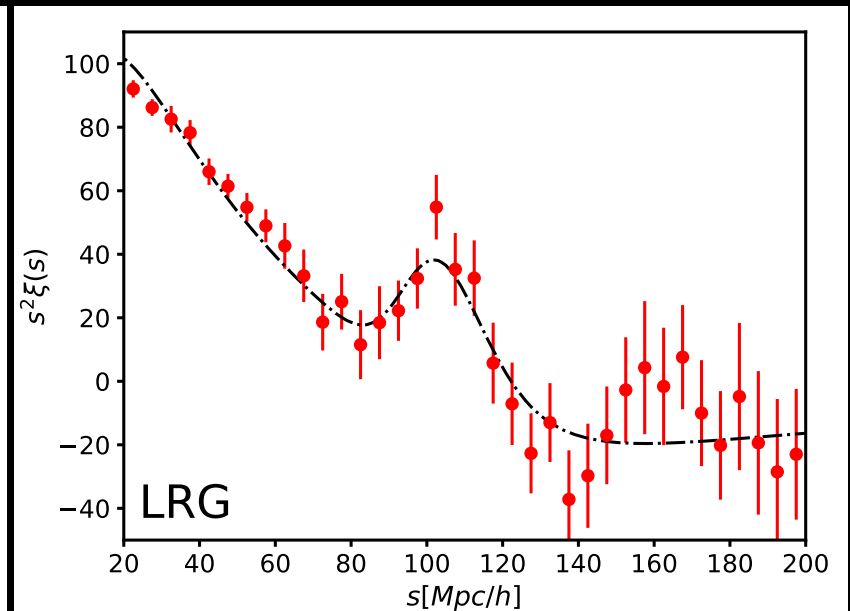
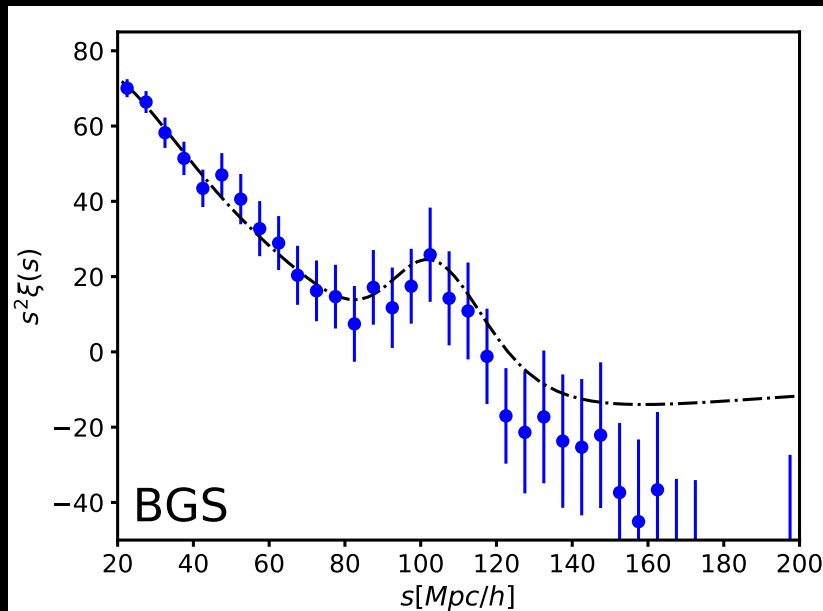
CORRELATION FUNCTION

linear prediction reasonable on large scales



GALAXY CORRELATION FUNCTION

- large scales: galaxies trace matter linearly



Moon et al. 2023 Early DESI data