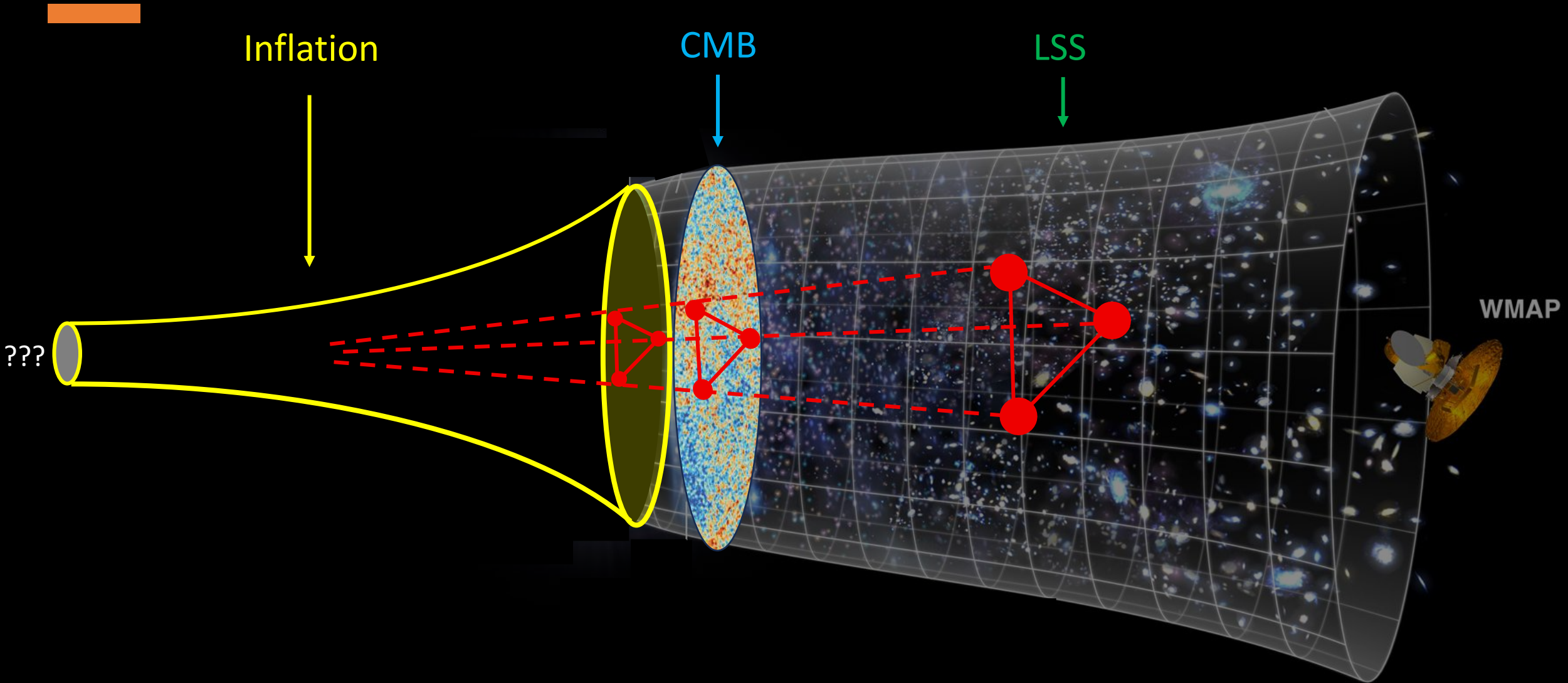


# Unitary renormalisation and the breaking of cosmological reality

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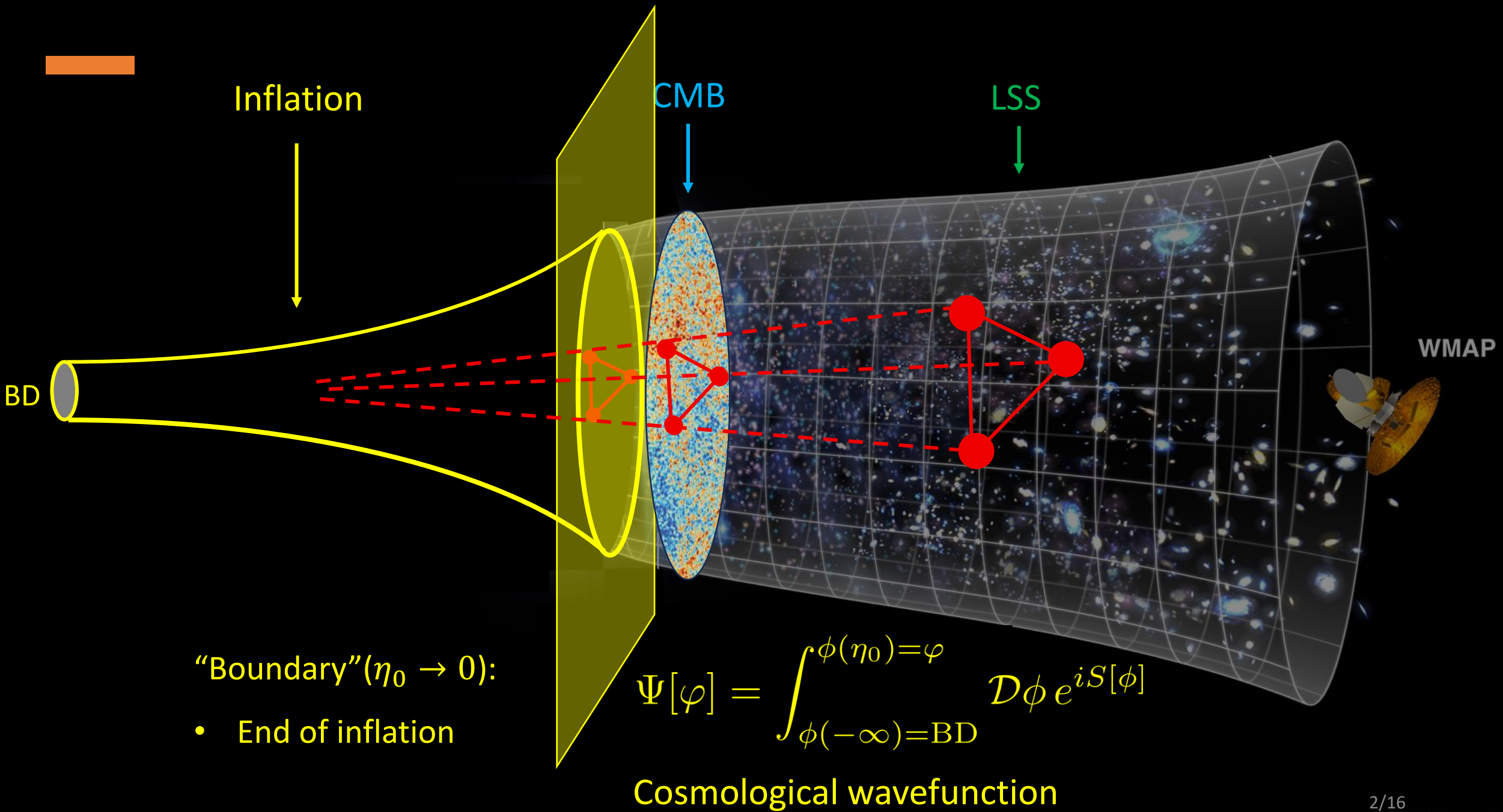
- Xi Tong
- DAMTP
- With Diksha Jain, Enrico Pajer,  
David Stefanyszyn, Yuhang Zhu,  
Tao Liu, Yi Wang and Zhong-Zhi Xianyu
- Based on
  - ◆ 2509.02696
  - ◆ 2309.07769 (JHEP)
  - ◆ 1909.01819 (JHEP)





- **Cosmological correlators** are a powerful probe of **inflationary physics** @  $H \leq 10^{13}$  GeV  
(DoFs, symmetries, fundamental principles, etc.)





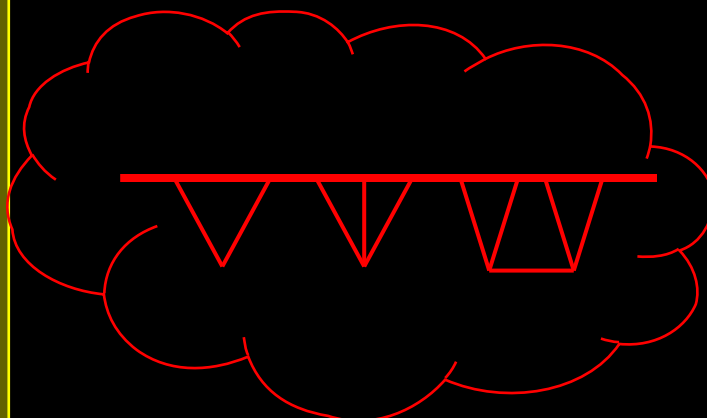
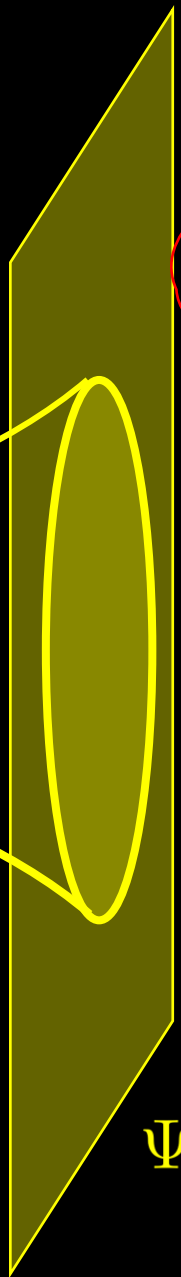
Inflation



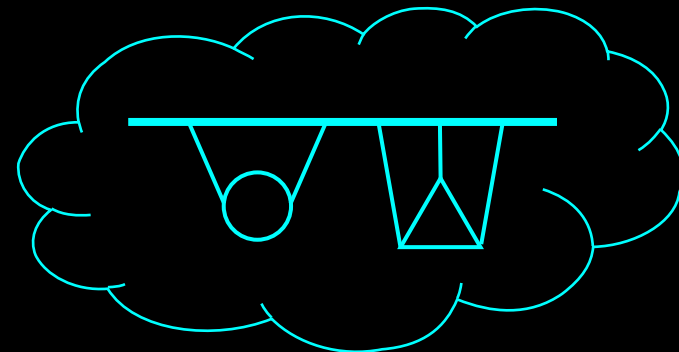
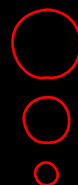
BD



- Path integral computed by **perturbative diagrammatics**



“Classical”  
tree diags



“Quantum”  
loop diags



=

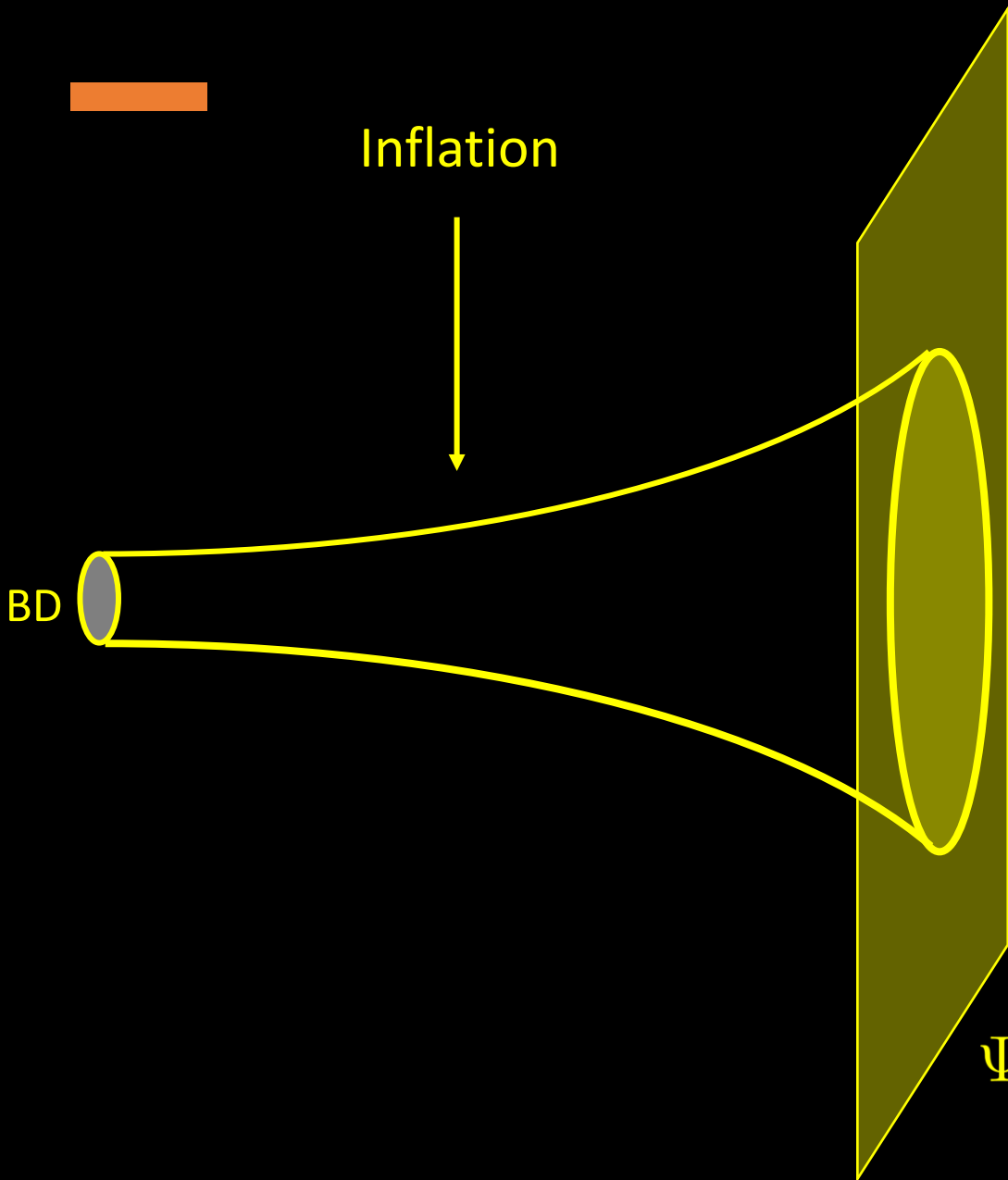
$$e^{iS[\phi_{cl}]}$$

×

$$\int \mathcal{D}\delta\phi e^{i(S[\phi_{cl}+\delta\phi]-S[\phi_{cl}])}$$

$$\Psi[\varphi] = \int_{\phi(-\infty)=BD}^{\phi(\eta_0)=\varphi} \mathcal{D}\phi e^{iS[\phi]}$$

Cosmological wavefunction



- Correlators are extracted from the WF coeffs via the **Born rule**

$$\langle \varphi(\mathbf{k}_1) \cdots \varphi(\mathbf{k}_n) \rangle = \frac{\int \mathcal{D}\varphi \left| \Psi[\varphi] \right|^2 \varphi(\mathbf{k}_1) \cdots \varphi(\mathbf{k}_n)}{\int \mathcal{D}\varphi \left| \Psi[\varphi] \right|^2}$$

$$\equiv \exp \left[ \sum_{n=2}^{\infty} \frac{1}{n!} \underbrace{\left( \psi_n^{0L} + \psi_n^{1L} + \psi_n^{2L} + \cdots \right)}_{\text{N-pt wavefunction coeffs}} \varphi^n \right]$$

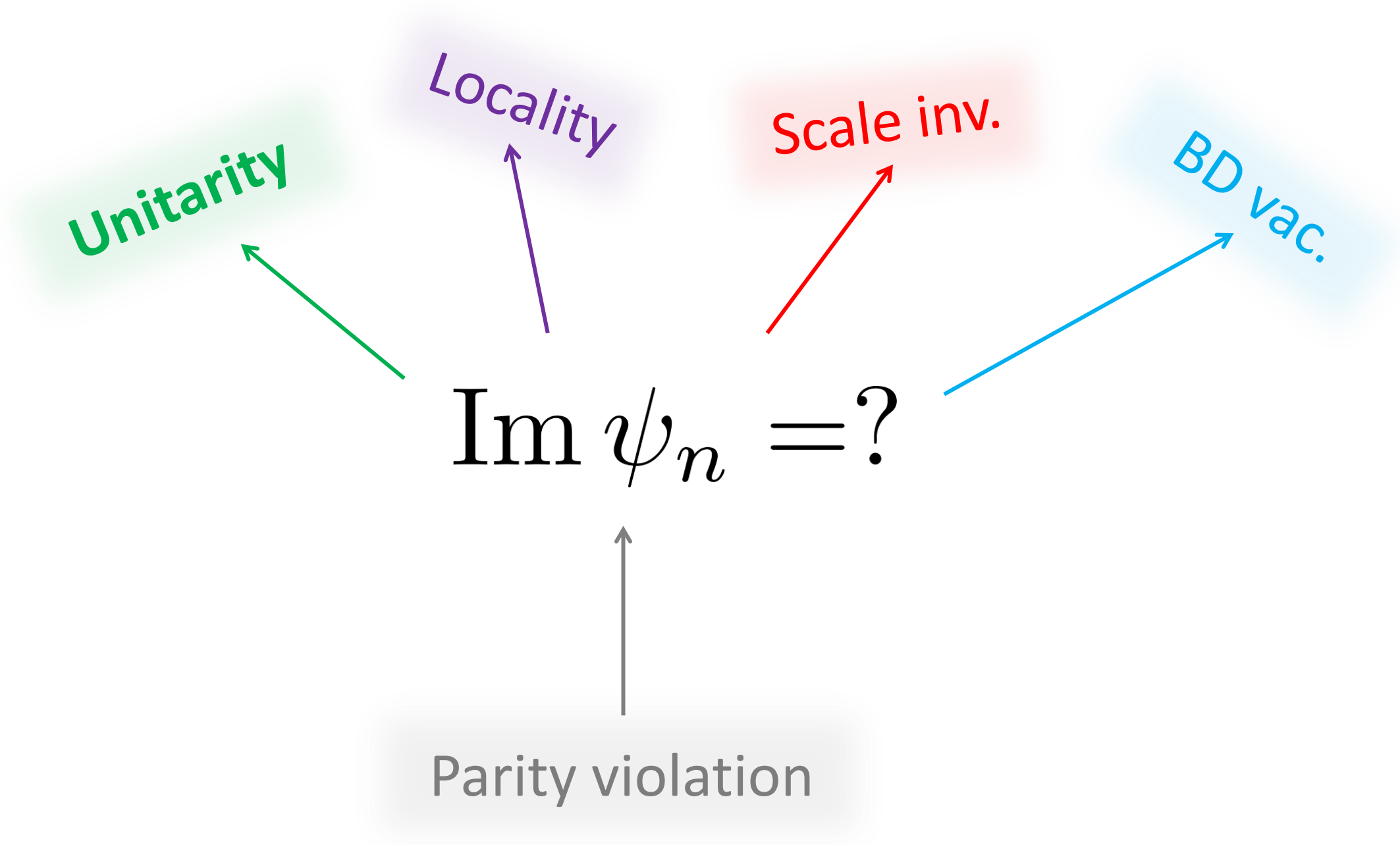
N-pt wavefunction coeffs

$$\Psi[\varphi] = \int_{\phi(-\infty)=BD}^{\phi(\eta_0)=\varphi} \mathcal{D}\phi e^{iS[\phi]}$$

Cosmological wavefunction are useful because...

... they make *fundamental principles* manifest

$$\operatorname{Im} \psi_n = ?$$





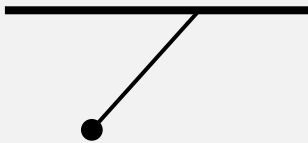
The tree-level reality:  $\text{Im } \psi_n^{0\text{L}} = 0$

[Liu, **Tong**, Wang & Xianyu, 2019]  
[Stefanyszyn, **Tong** & Zhu, 2023]

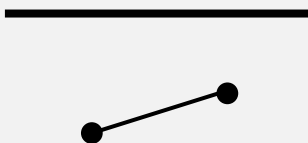
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[Liu, **Tong**, Wang & Xianyu, 2019]  
[Stefanyszyn, **Tong** & Zhu, 2023]

- Feynman rules:



$$= K_k(\eta) = (1 - ik\eta)e^{ik\eta}$$

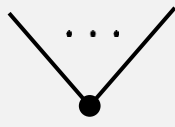


$$= G_k(\eta_1, \eta_2)$$

$$= \frac{iH^2}{k^3} (k\eta_1 \cos k\eta_1 - \sin k\eta_1)$$

$$\times (1 - ik\eta_2)e^{ik\eta_2} \theta(\eta_1 - \eta_2)$$

$$+ (\eta_1 \leftrightarrow \eta_2)$$

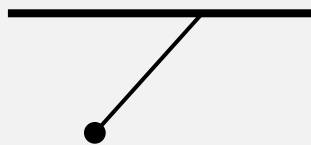


$$= i\lambda \int_{-\infty(1-i\epsilon)}^0 \frac{d\eta}{(-H\eta)^4} \mathcal{L}_{\text{int}}(\eta \partial_\eta, -i\mathbf{k}\eta)$$

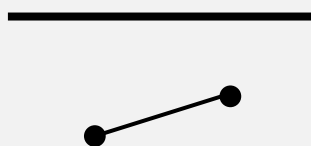
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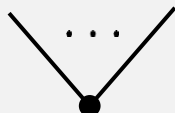


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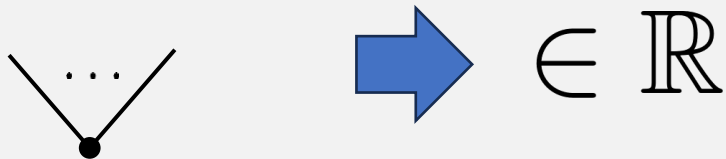
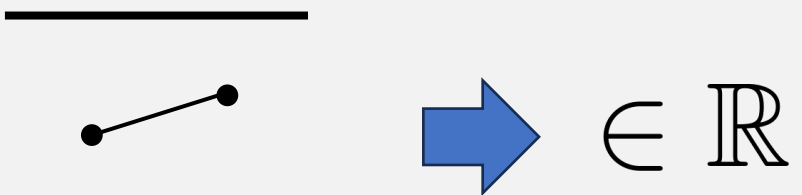
$$= i\lambda \int_{-\infty(1-i\epsilon)}^0 \frac{d\eta}{(-H\eta)^4} \mathcal{L}_{\text{int}}(\eta \partial_\eta, -i\mathbf{k}\eta)$$

- $i$  &  $\eta$  always appear together  
(BD vac.)
- $\partial_\eta$  &  $\mathbf{k}$  always appear with  $\eta$   
(scale inv.)
- $\lambda$  's are real  
(unitarity)

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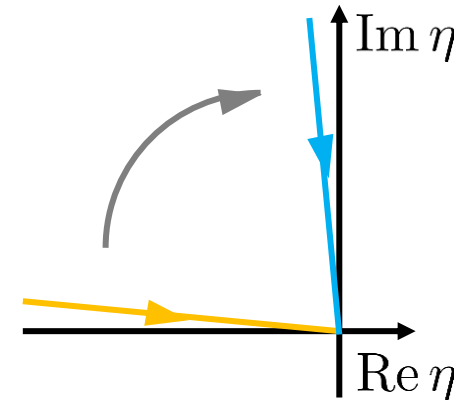
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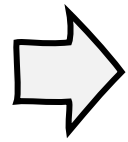
- Wick rotation  $\eta = i\chi$   
(locality)



Now, what does this cosmological reality imply?

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$$\varphi^*(\mathbf{k}) = \varphi(-\mathbf{k})$$



Real = Parity-even!



# A no-go theorem on parity violation

- Unitarity
- locality
- Scale inv.
- BD vac.
- Tree

No PV correlators  
 $\langle \varphi(\mathbf{k}_1) \cdots \varphi(\mathbf{k}_n) \rangle^{\text{PO}} = 0$   
in massless scalar EFTs

[Liu, Tong, Wang & Xianyu, 2019]

[Cabass, Jazajeri, Pajer & Stefanyszyn, 2022]

Generalisable to

- Higher spins
- Other dimensions

[Cabass, Jazajeri, Pajer & Stefanyszyn, 2022]

[Goodhew, Thavanesan & Wall, 2024]

[Thavanesan, 2025]

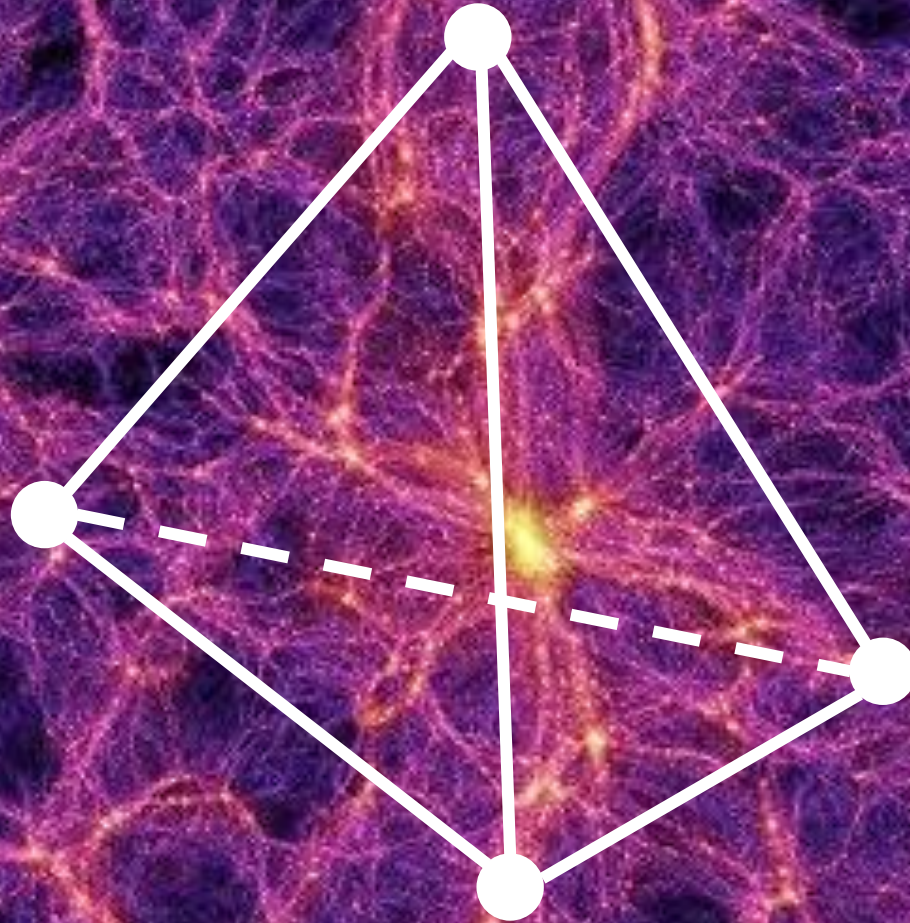
➤ An infinite & invisible tower of PV operators:

$$\mathcal{L}_{\text{int}} = \frac{g}{2} \ddot{\phi}^2 + \cdots + \frac{\lambda}{3!} \dot{\phi}^3 + \cdots + c_{10} \epsilon_{ijk} \phi \partial_i \phi \partial_j \dot{\phi} \partial_k \partial^2 \phi + \cdots$$

**dynamical** but **invisible**

➤ A **null test** on the **fundamental principles**





[See Zucheng Gao's talk]

$$\langle \zeta^4 \rangle \sim \text{Galaxy}^4$$

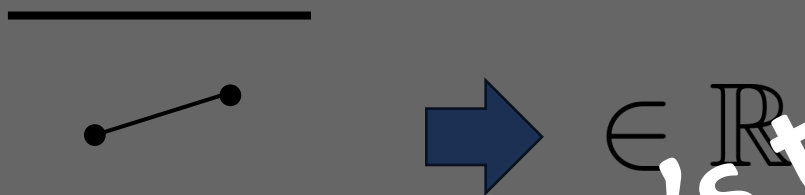


✓ Cosmo PV is growingly active.

◆ But how about loops??

The tree-level story:  $\text{Im } \psi_n^{0L} = 0$  [Liu, Tong, Wang & Xianyu, 2019]  
[Stefanyszyn, Tong & Zhu, 2023]

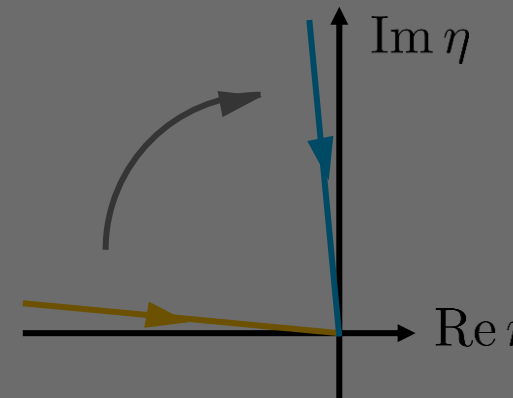
- Feynman rules:



Where's the "tree" assumption?

- $i$  &  $\eta$  always appear together  
(BD v.c. + locality)
- $\partial_\eta$  &  $\mathbf{k}$  always appear with  $\eta$   
(scale inv.)
- $\lambda$  's are real  
(unitarity)

➤ Wick rotation  $\eta = i\chi$



# How about loops?

Manifestly real after Wick rot.  $\eta = i\chi$

- At 1 loop:

$$\psi_n^{1L} = \int \frac{d^3\mathbf{q}}{(2\pi)^3} \int_{-\infty}^0 \overbrace{\left[ \prod_{v=1}^V i d\eta_v f_v(\eta \partial_\eta, -i\mathbf{k}\eta) \right] \left[ \prod_{e=1}^n K_e(\eta_e, k_e) \right] \left[ \prod_{i=1}^I G_e(\eta_i, \eta'_i, q_i) \right]} = \text{Real?}$$

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Careful: UV div. !



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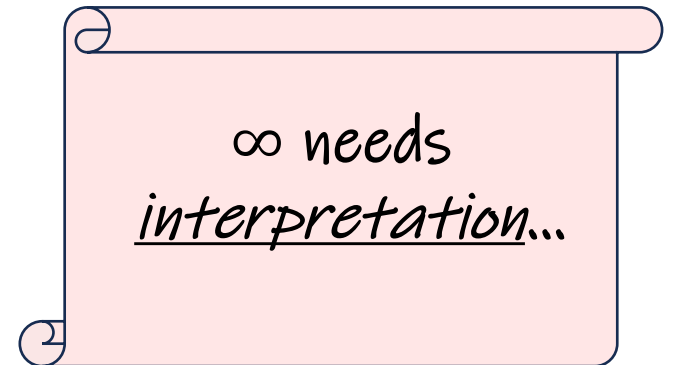
Careful: UV div. !

- A divergent integral of real numbers may not be real anymore...

$$(+\infty)^* - (+\infty) = 0?$$

- Classic example:

$$1 + 2 + 3 + \dots \rightarrow -\frac{1}{12}$$

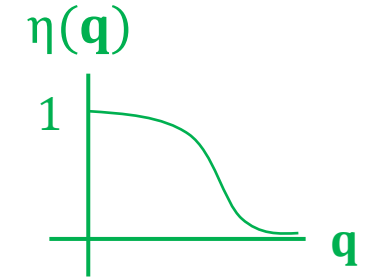


# The breaking of cosmological reality

- ( $\eta$ -) regularisation:

[Padilla & Smith, 2024,2024]  
(Minkowski)

$$\int \frac{d^3\mathbf{q}}{(2\pi)^3} \rightarrow \int \frac{d^3\mathbf{q}}{(2\pi)^3} \times \eta(\mathbf{q})$$



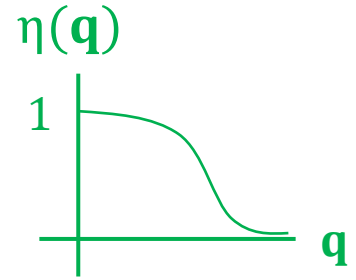
**Not** necessarily real!

# The breaking of cosmological reality

- ( $\eta$ -) regularisation:

[Padilla & Smith, 2024,2024]  
(Minkowski)

$$\int \frac{d^3\mathbf{q}}{(2\pi)^3} \rightarrow \int \frac{d^3\mathbf{q}}{(2\pi)^3} \times \eta(\mathbf{q})$$



**Not necessarily real!**

- Not any  $\eta$  goes in dS/inflation...

**Unitary & analytic  $\eta$ -regs**

Unitarity

Scale inv.

Causality



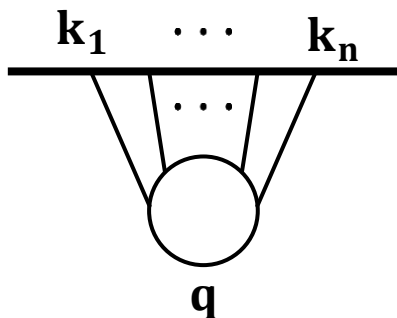
$$\eta = \rho \left( \frac{q}{ik} / \frac{\Lambda}{H} \right)$$

$$\rho(0) = 1 ,$$

$$\rho(\infty e^{i\alpha}) = 0 , \quad -\frac{\pi}{2} < \alpha < \frac{\pi}{2}$$

# A 1-loop universality

Real by tree-level reality



$$\begin{aligned}
 &= \psi_n^{1L} = \int_{\min}^{\infty} \frac{dq_+}{q_+} f(\{k\}, \{q\}, \{\mathbf{k}\}) \times \rho \left( \frac{q_+/k_T}{i\Lambda/H} \right) + \text{counterterms} \\
 &= \sum_{m=\text{even}} f_m C_m[\rho] \left( \frac{\Lambda}{H} \right)^m + f_0 \left( \log \frac{\Lambda}{H} + \frac{i\pi}{2} + \gamma[\rho] + g_0 \right) + \dots
 \end{aligned}$$

- A **universal** imaginary part for all 1-loop renormalised WFs:

$$\left( \mu \frac{\partial}{\partial \mu} - \frac{2}{\pi} \text{Im} \right) \hat{\psi}_n^{1L} = 0$$

- Any 1-loop topologies
- Any bulk fields with **integer spin** and **light mass**
- Any **IR-conv.** interactions
- Any U & A  $\eta$ -regs (**uncountably** many)

# Summary

---

- ✓ Cosmological WF and correlators are useful
- ✓ WFs satisfy reality at tree level
- ✓ Translates to a no-go thrm. for PV with obs. consequences
- ✓ Yet spontaneously broken by UV divs loops
- ✓ By U & A, breaking is universal and hints at a connection to RG in dS

$$\left( \mu \frac{\partial}{\partial \mu} - \frac{2}{\pi} \text{Im} \right) \hat{\psi}_n^{1L} = 0$$



Thanks for  
your attention!





# Outlooks

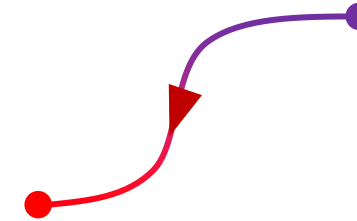
□ Connection to RG? — — — — — →

□ **Parity violation as a “scale anomaly”?** — — — — — →

□ More on  $\eta$  reg in cosmology — — — — — →

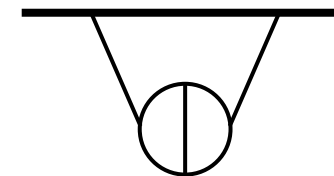
□ Fermions and higher loops? — — — — — →

□ Etc.



$$\langle \phi^n \rangle_{\text{PO}} = \frac{i\pi}{2} \mu \frac{\partial}{\partial \mu} \langle \phi^{n-1} \pi \rangle_{\text{PO}} ?$$

$$\eta(k_{\text{IR}}) ?$$



Back ups

# A minimal set up

- Single-field slow-roll inflation

$$S = \int d^4x \sqrt{-g} \left[ \frac{1}{2} M_p^2 R - \frac{1}{2} (\partial\phi)^2 - V(\phi) + \mathcal{L}_{\text{int}} \right]$$

- Quasi-de Sitter (dS) background

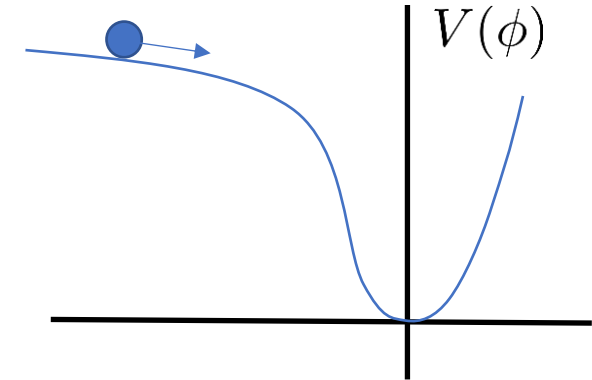
$$ds^2 = a^2(\eta) (-d\eta^2 + d\mathbf{x}^2) \quad \text{with} \quad a(t) = e^{Ht} = -\frac{1}{H\eta}$$

- A single massless scalar DoF with (**IR convergent**) self-interactions

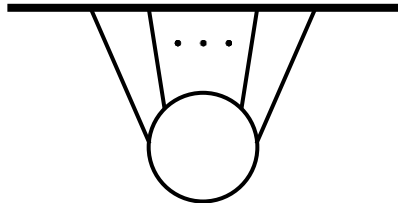
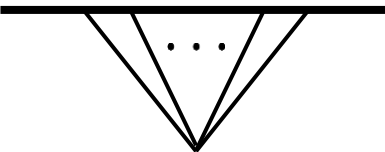





$$\mathcal{L}_{\text{int}} = \mathcal{L}_{\text{int}}(\partial\phi, \partial^2\phi, \dots) = \frac{\lambda}{3!} \dot{\phi}^3 + \dots$$

(Isolate the **sub-horizon** physics)

Shift symmetric



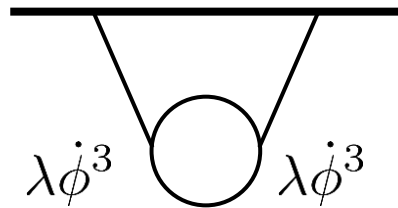
# Checking with dim regs

Imaginary part	 $\psi_n^{1L}$  $\psi_n^{ct}$ 
dim reg ( $d = 3 - \epsilon, m_\phi^2 = 0$ )	$0$  $\frac{i\pi}{2}$  $\text{Im } \hat{\psi}_n^{1L} = \frac{\pi}{2}$
mass-dim reg ( $d = 3 - \epsilon, m_\phi^2 = -\frac{3}{2}\epsilon H^2$ )	$\frac{i\pi n}{4}$  $-\frac{i\pi(n-2)}{4}$  $\text{Im } \hat{\psi}_n^{1L} = \frac{\pi}{2}$

# Counterterms can't solve the issue

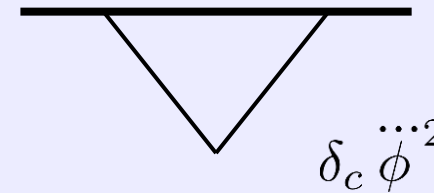
- Renormalisation

$$\hat{\psi}_2^{1L} = \psi_2^{1L} + \psi_2^{ct}$$



$$\frac{1}{15} \frac{\lambda^2 H^2}{(4\pi)^2} k^3 \left( \frac{1}{\epsilon} + \frac{i\pi}{2} + \text{real \& fin.} \right)$$

$$\frac{1}{15} \frac{\lambda^2 H^2}{(4\pi)^2} k^3 \left( \log \frac{\Lambda}{H} + i\theta + \text{real \& fin.} \right)$$



$$\frac{1}{15} \frac{\lambda^2 H^2}{(4\pi)^2} k^3 \left( -\frac{1}{\epsilon} \right)$$

$$\frac{1}{15} \frac{\lambda^2 H^2}{(4\pi)^2} k^3 \left( -\log \frac{\Lambda}{H} \right)$$

Real by the tree-level reality



Hermitian counterterms **cannot** alter the imaginary part