The background of the slide is a composite image. The left side features a portion of a Cosmic Microwave Background (CMB) fluctuation map, showing a granular pattern of blue and orange colors. The right side shows a deep-field image of a galaxy cluster, with numerous galaxies of various shapes and colors (red, blue, white) against a dark background, some with prominent blue star trails. Two horizontal white lines are positioned above and below the main title.

Positive Geometry in Cosmological Correlators

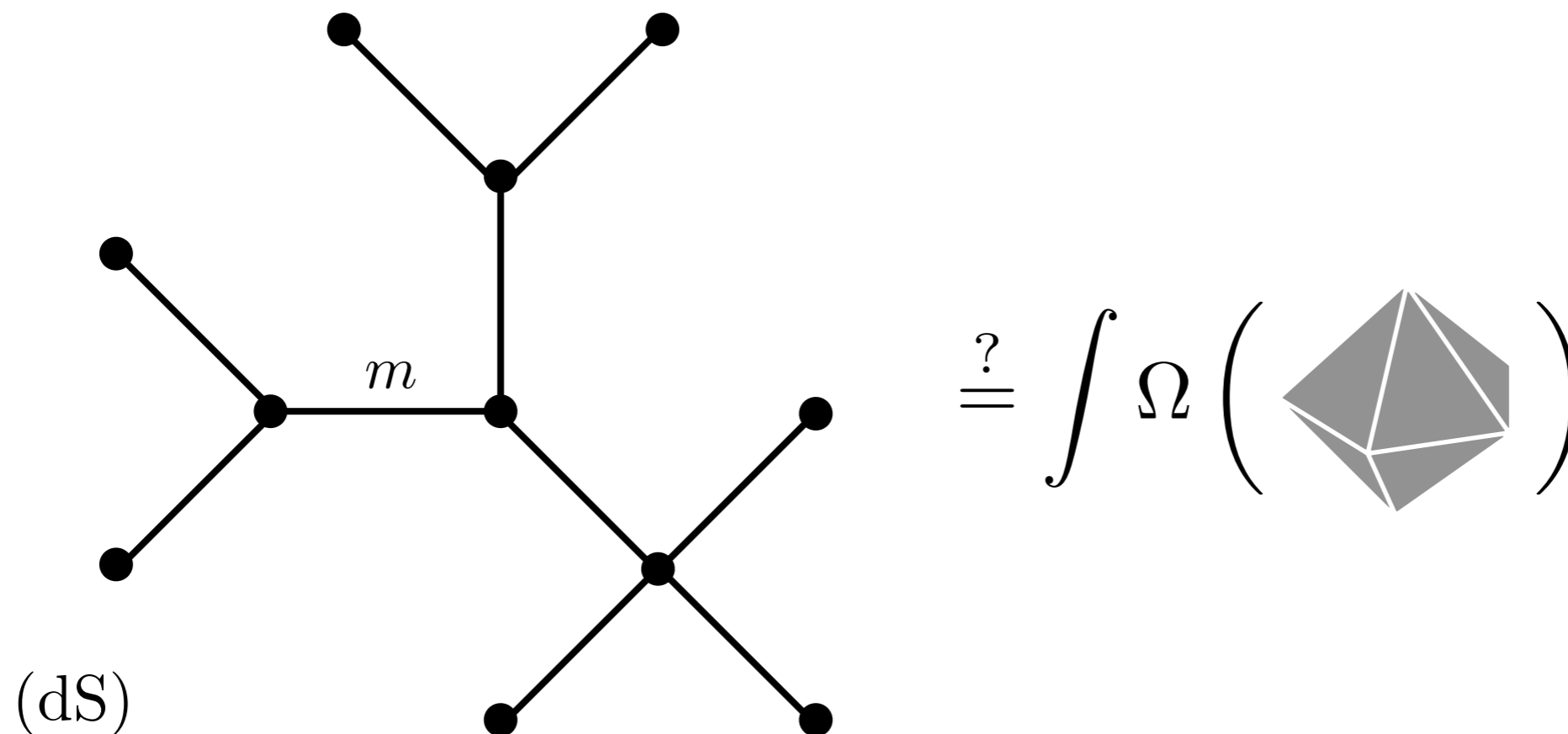
Denis Werth

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UNIVERSE+

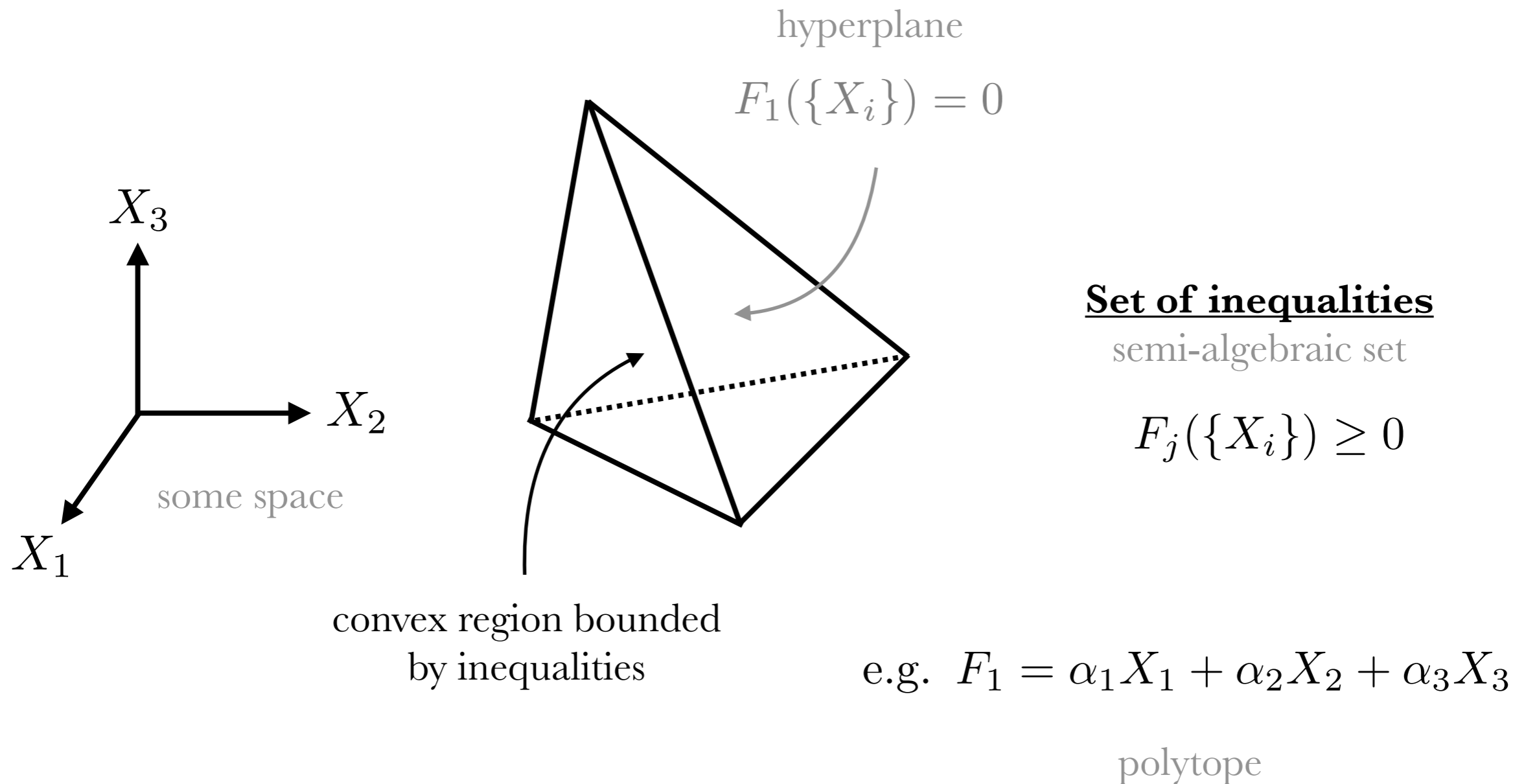
Can massive de Sitter cosmological correlators
be described by positive geometries?



Short answer: We don't know yet!

Goal of the talk: Give you the tools to understand the question & why it's worth asking

What is a *positive geometry* in simple terms?

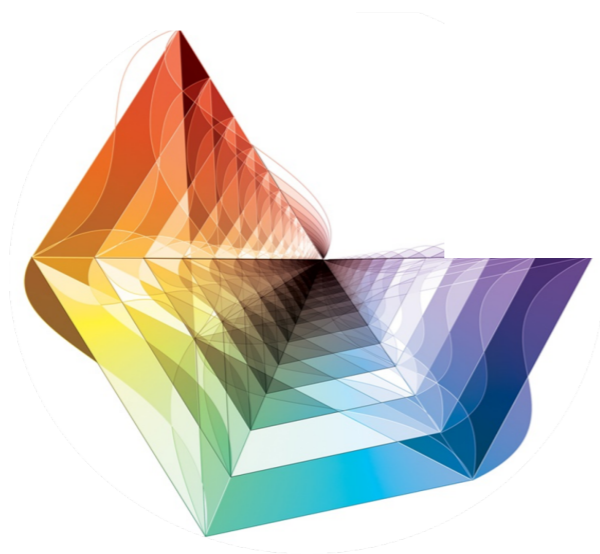


Positive geometry is a *new emergent field of mathematics* that is related to and inspired by physics

Positive geometries can describe (flat-space)
scattering amplitudes in *some* theories

Amplituhedron

(Planar $N = 4$ SYM)
tree level & loop integrands

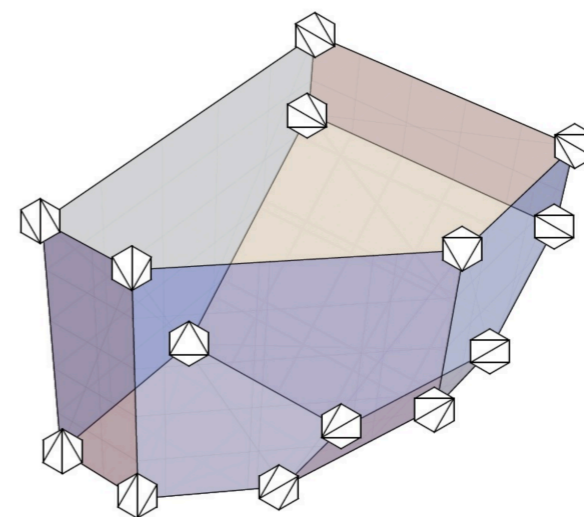


[Arkani-Hamed, Trnka '13]

momentum twistor space/
positive Grassmanian

ABHY Associahedron

($\text{Tr}(\phi^3)$ / bi-adjoint ϕ^3)



[Arkani-Hamed, Bai, He, Yan '17]

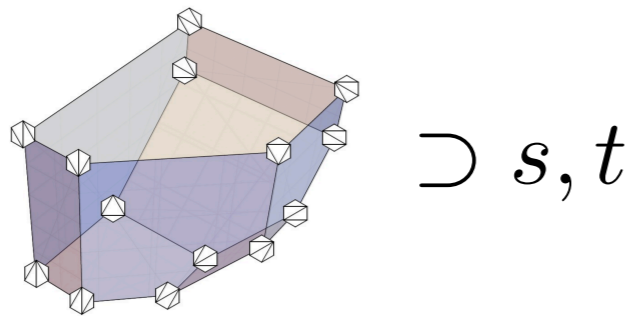
kinematic space of Mandelstam
invariants

Surfacehedron generalisation to loop integrands

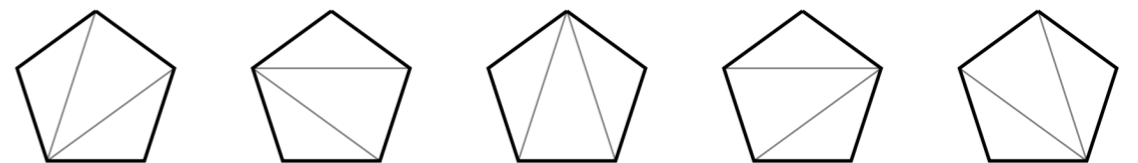
[Arkani-Hamed, Frost, Salvatori, Plamondon, Thomas '23]

Novel properties emerge from positive geometries

No Feynman diagrams

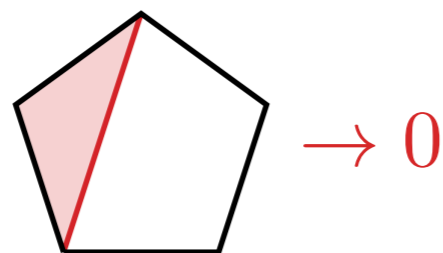


New decompositions
from triangulations

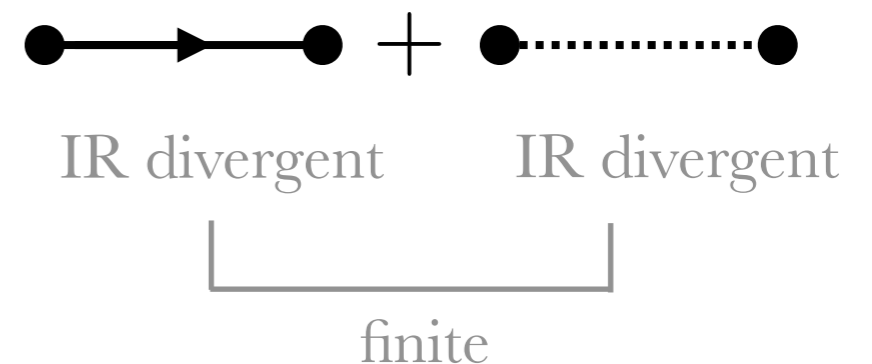


Physical properties made manifest

e.g. singularities encoded in the facet structure of the positive geometry

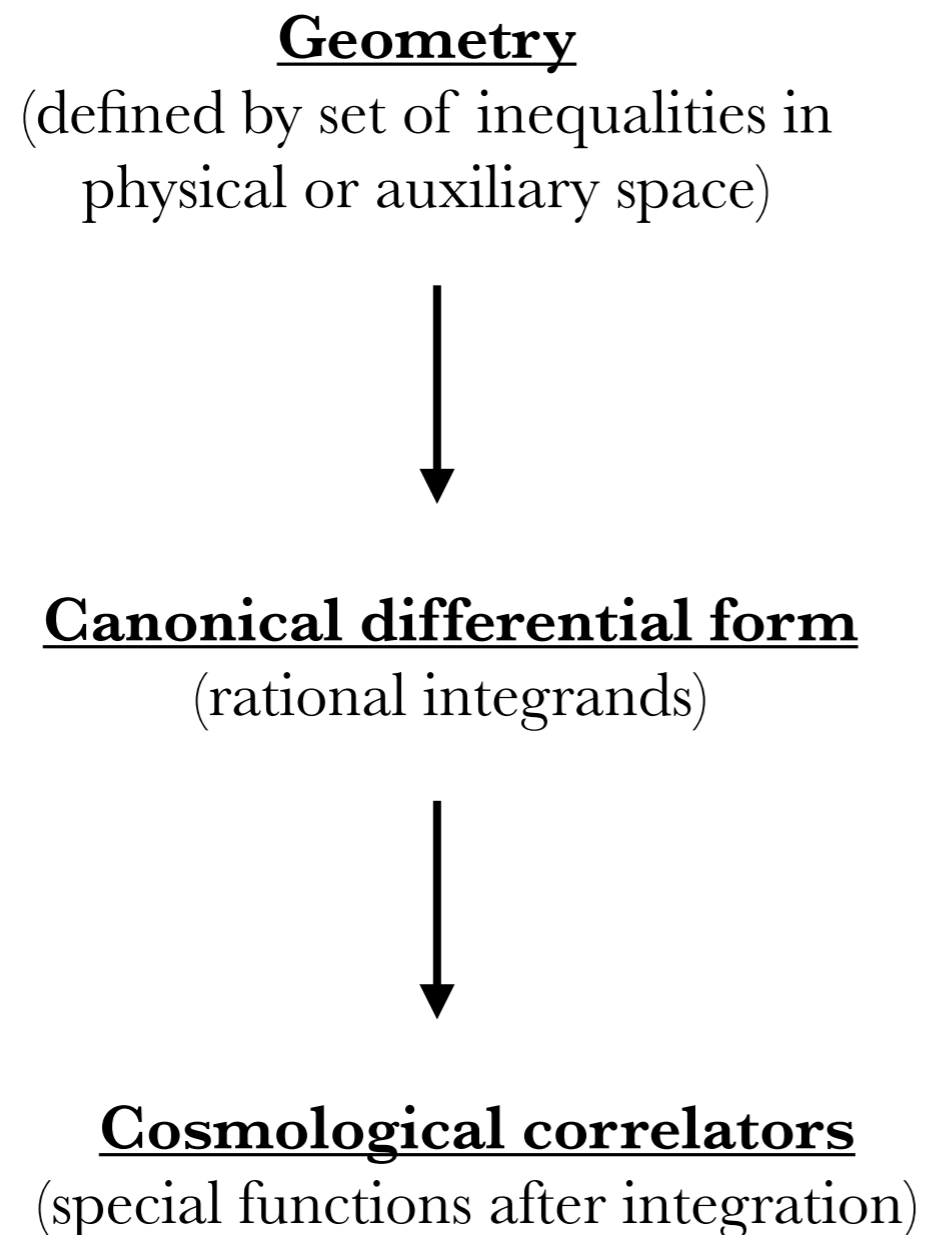


No spurious divergences



No computational advantage but rather new conceptual viewpoint!

Positive geometry can bring a *novel fundamental perspective* on cosmological correlators



What geometries are
relevant?



(?)

How to construct the
forms?

$$\Omega \left(\text{polyhedron} \right) \quad (?)$$

How to perform the
integration?

$$\int \Omega \left(\text{polyhedron} \right) \quad (?)$$

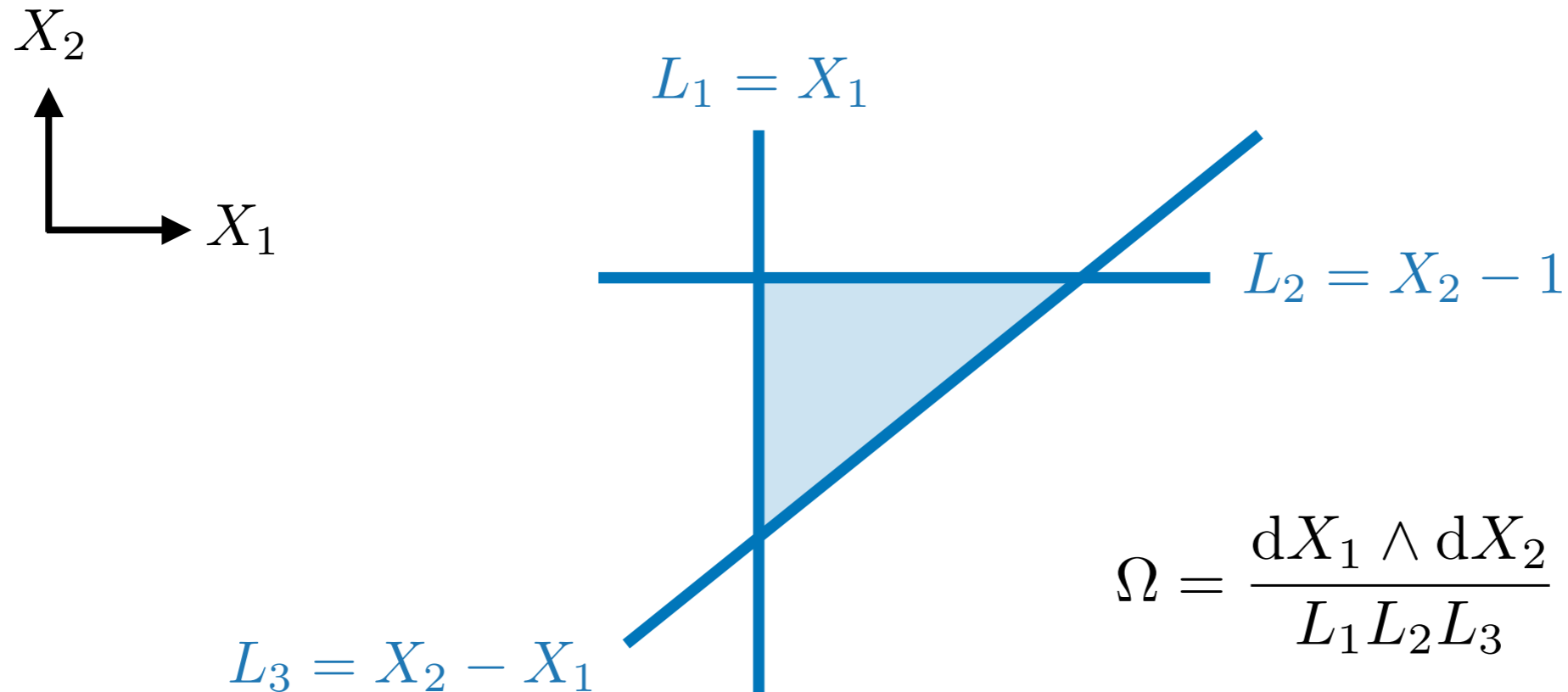
Interval (1D)

The canonical differential form should have logarithmic singularities at the endpoints $\begin{cases} X + a = 0 \\ X + b = 0 \end{cases}$



$$\Omega = \frac{dX}{X + a} - \frac{dX}{X + b} = \frac{(b - a)dX}{(X + a)(X + b)} = d \log \left(\frac{X + a}{X + b} \right)$$

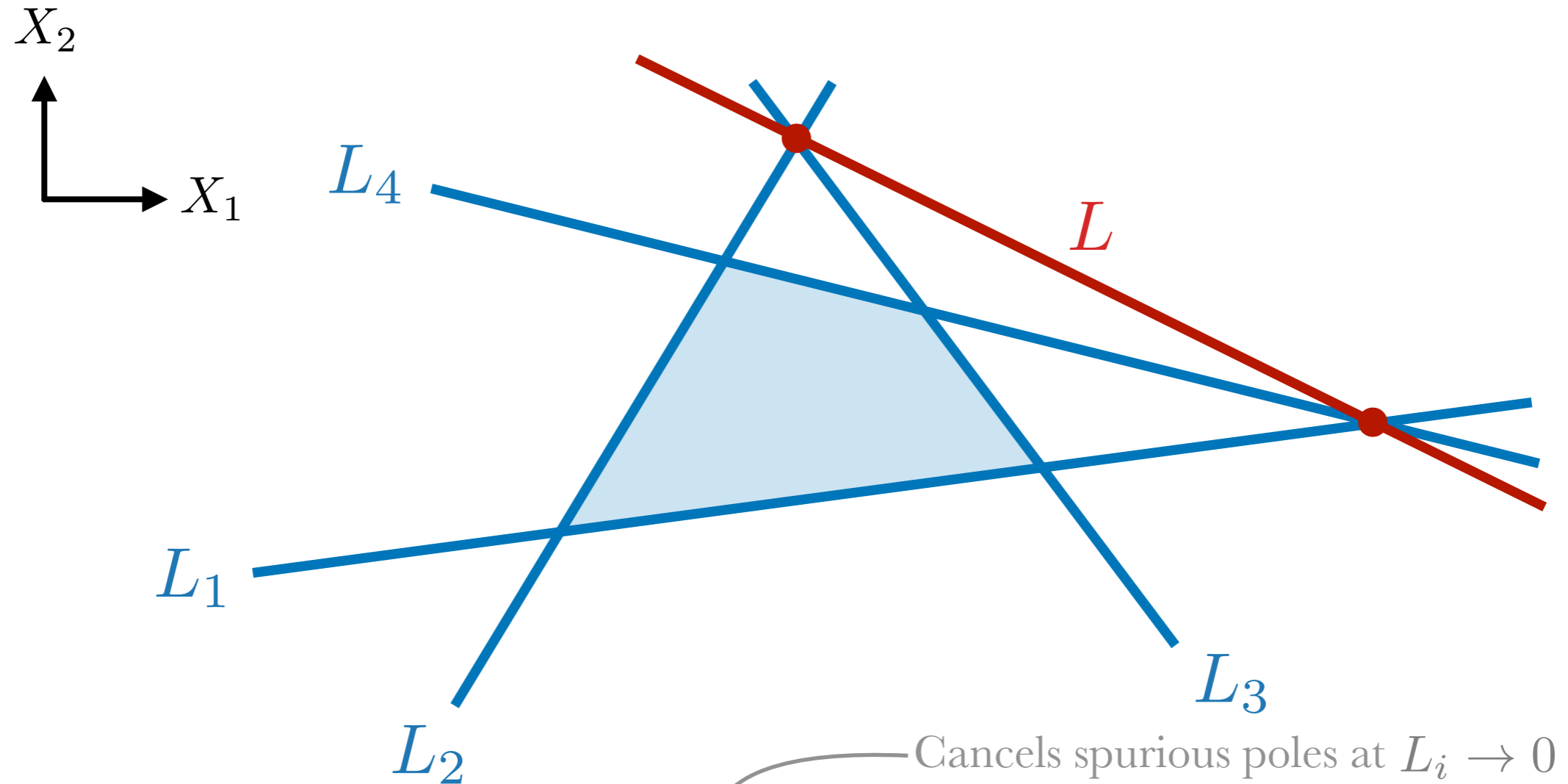
Triangle (2D)



$$\text{Res}_{L_2 \rightarrow 0} [\Omega] = \frac{dX_1}{X_1(1 - X_1)} \quad \longleftarrow \text{canonical form of the interval}$$

Definition: Unique differential form with logarithmic singularities (only) on all the boundaries of the space, with residues on each boundary given by the canonical form on that boundary

Quadrilateral (2D)



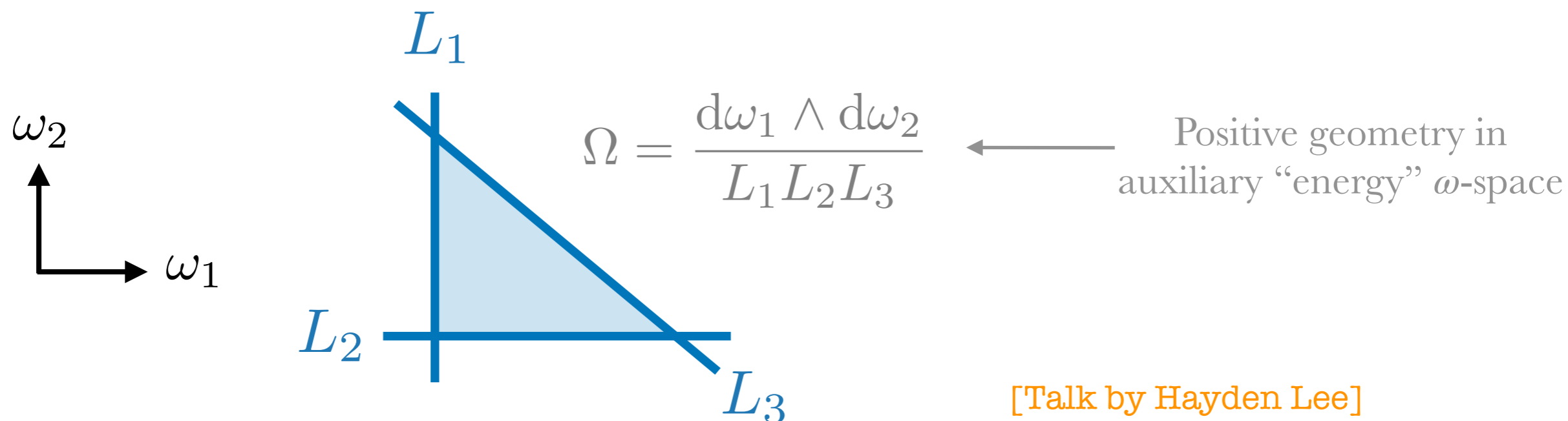
$$\Omega = \frac{L \, dX_1 \wedge dX_2}{L_1 L_2 L_3 L_4}$$

This recursive definition & construction of canonical forms generalises to higher dimensions

De Sitter correlators of a *self-interacting conformally coupled scalar field* emerge from positive geometries

$$S = \int d^4x \sqrt{-g} \left[-\frac{1}{2} (\partial_\mu \phi)^2 - \frac{R}{12} \phi^2 - \sum_n \frac{\lambda_n}{n!} \phi^n \right]$$

$$\begin{array}{c} X_1 \quad X_2 \\ \bullet \text{---} \bullet \\ Y \end{array} = \int_0^\infty \frac{d\omega_1 \wedge d\omega_2}{(X_1 + \omega_1 + Y)(X_2 + \omega_2 + Y)(X_1 + \omega_1 + X_2 + \omega_2)}$$



This generalises to arbitrary tree graphs and loop integrands

Massive de Sitter correlators have integral representations
with *hypergeometric kernels* & *rational integrands*

$$S = \int d^4x \sqrt{-g} \left[-\frac{1}{2} (\partial_\mu \phi)^2 - \frac{R}{12} \phi^2 - \frac{1}{2} (\partial_\mu \sigma)^2 - \frac{1}{2} m^2 \sigma^2 - \sum_{m,n} \frac{\lambda_{m,n}}{n!m!} \phi^m \sigma^n \right]$$

Legendre function
(encodes the mass)

rational

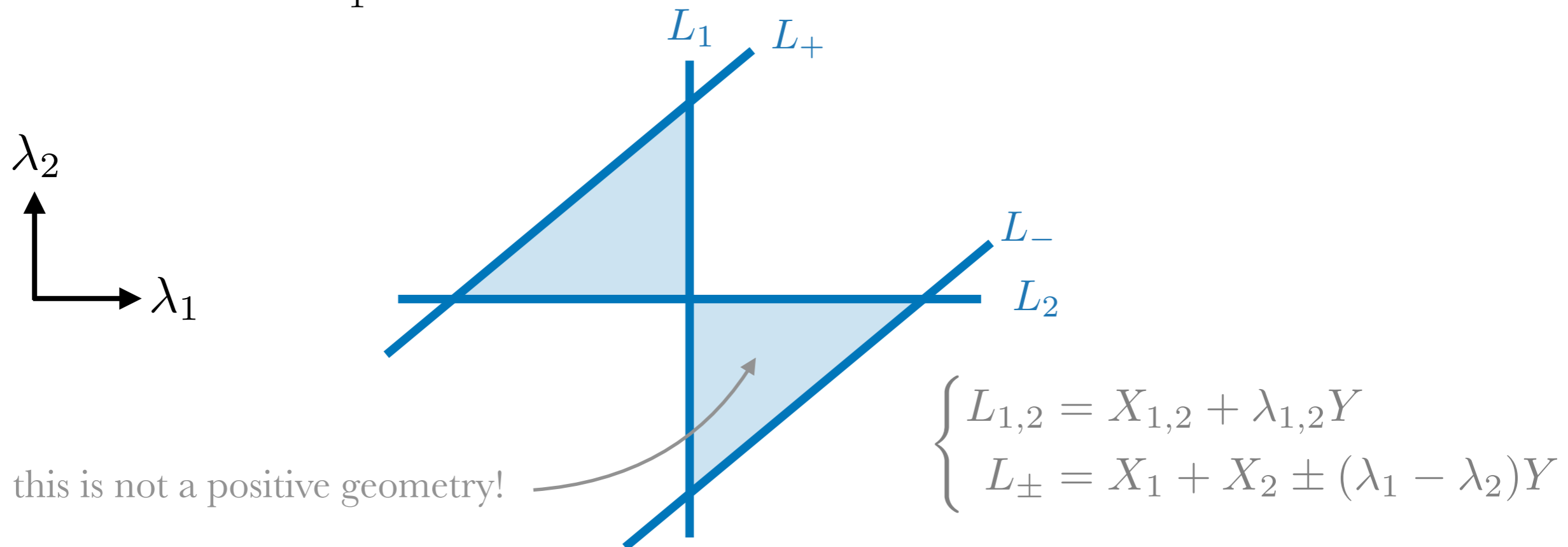
$$\begin{array}{c} X_1 \quad X_2 \\ \bullet \text{---} \bullet \\ Y \end{array} = \int_1^\infty P_{i\mu}(\lambda_1) P_{i\mu}(\lambda_2) d\lambda_1 d\lambda_2 \left(\frac{1}{[X_1 + X_2 + (\lambda_1 - \lambda_2)Y][X_1 + \lambda_1 Y]} \right. \\ \left. + \frac{1}{[X_1 + X_2 + (\lambda_2 - \lambda_1)Y][X_2 + \lambda_2 Y]} - \frac{1}{[X_1 + \lambda_1 Y][X_2 + \lambda_2 Y]} \right)$$

[Talk by Nathan Belrhali]

This generalises to arbitrary tree graphs

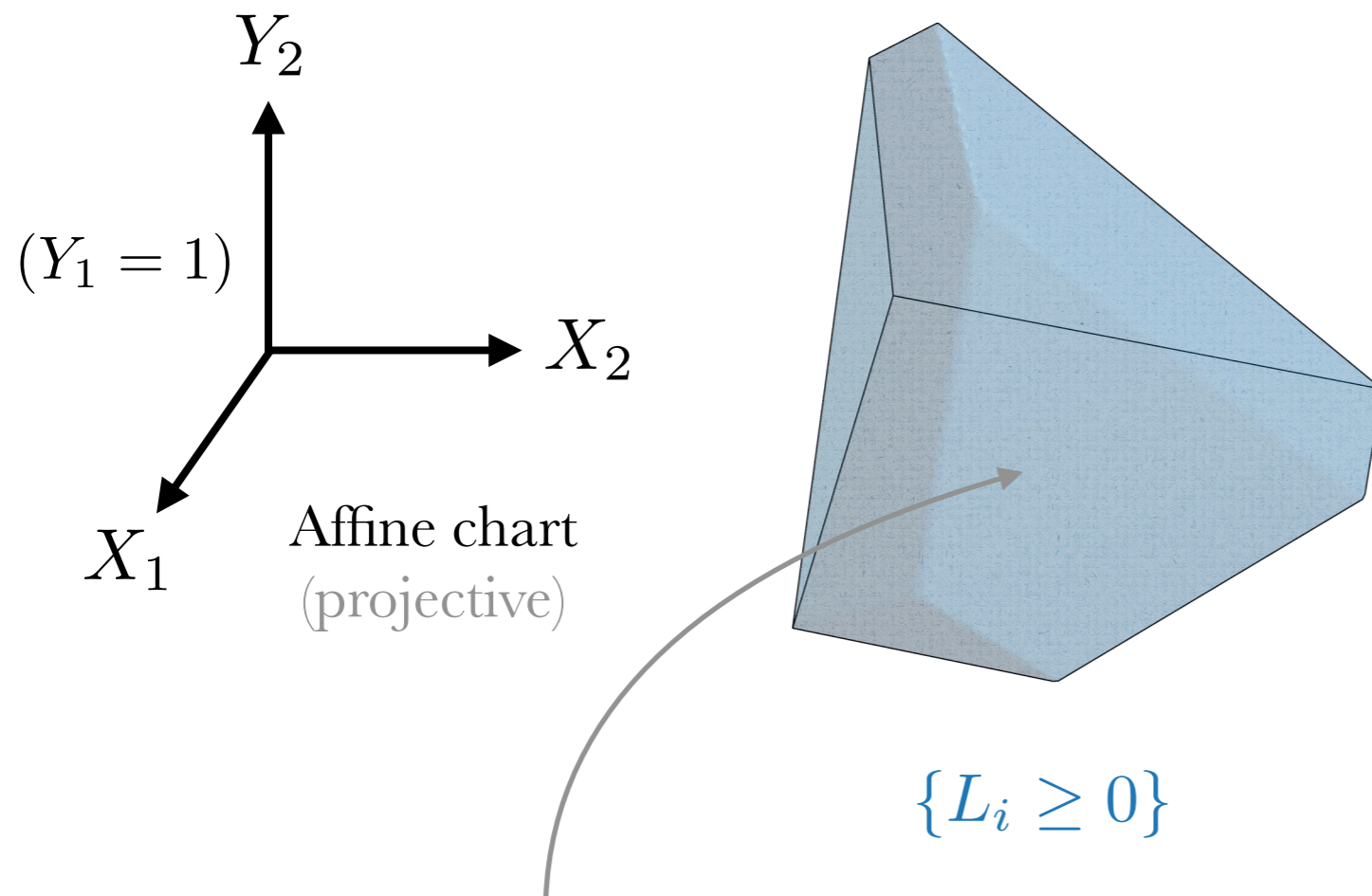
... but the rational integrands *are not* canonical
differential forms of positive geometries in the λ -space...

$$\begin{array}{c} X_1 \qquad X_2 \\ \bullet \text{---} \bullet \\ Y \end{array} = \int_1^\infty P_{i\mu}(\lambda_1) P_{i\mu}(\lambda_2) d\lambda_1 d\lambda_2 \left(\frac{1}{L_1 L_+} + \frac{1}{L_2 L_-} - \frac{1}{L_1 L_2} \right)$$



$$\Omega \stackrel{!}{=} \frac{d\lambda_1 \wedge d\lambda_2}{L_1 L_2 L_+} + \frac{d\lambda_1 \wedge d\lambda_2}{L_1 L_2 L_-} = \frac{(L_+ + L_-) d\lambda_1 \wedge d\lambda_2}{L_1 L_2 L_+ L_-}$$

... however we obtain a positive geometry in the
projective kinematic space



$$(Y_{1,2} = \lambda_{1,2}Y)$$

$$\begin{cases} L_{1,2} = X_{1,2} + Y_{1,2} \\ L_{\pm} = X_1 + X_2 \pm (Y_1 - Y_2) \end{cases}$$

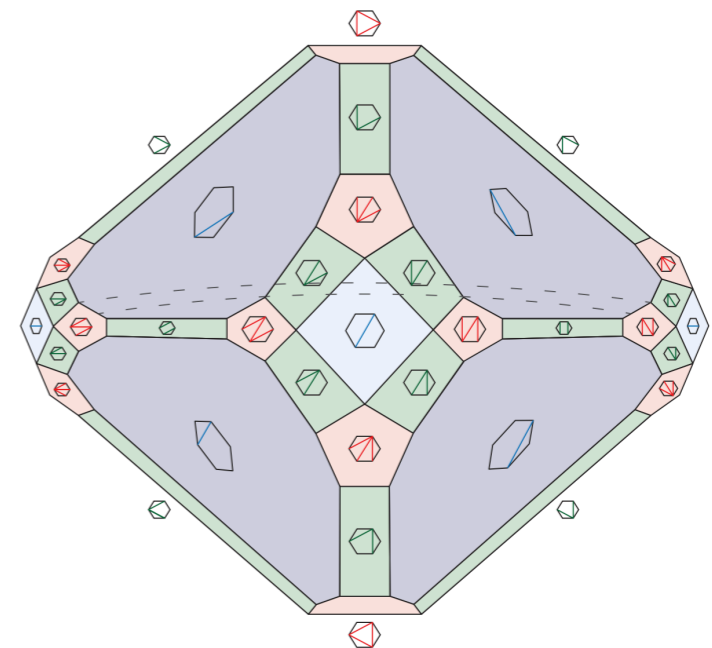
we need a slight deformation of the definition to interpret the integrand
as the canonical differential form associated to this positive geometry

Work in progress!

Many open questions...

- Is there a **simple & systematic canonical form (deformation)** for massive dS correlators?
- Can we read off the **complexity** (transcendental weight) from a suitable integral representation of massive dS correlators?
- How to evaluate the integrals? (Landau analysis, canonical differential equations, ...)
- **Can we go beyond individual graphs?** (towards dS massive cosmohedron)
- ...

$$\Omega \rightarrow \tilde{\Omega}$$



Work in progress!

The background of the slide is a large-scale visualization of the cosmic web, showing a complex network of galaxy clusters and filaments in shades of blue and orange. In the bottom-left corner, there is a circular inset showing a noisy, blue and orange pattern, likely representing a simulation or data visualization of the early universe. In the top-right corner, there is a circular inset showing a dense field of galaxies, with some highlighted by blue lines, possibly representing a specific region of interest or a different type of astronomical data.

Thank you